## Physics of plasma mirrors in ultraintense laser fields

Fabien Quéré <u>ELI-ALPS</u> & PASQAL formerly CEA-Saclay Naturally (or almost so) produced on initially solid targets by <u>intense ultrashort</u> laser pulses



## Why studying plasma mirrors?

Naturally (or almost so) produced on initially solid targets by <u>intense ultrashort</u> laser pulses



#### Plasma frequency

$$\omega_p^2(x) = n_e e^2 / m \mathcal{E}_0 >> \omega_L^2$$

with  $n_e \approx 10^{23} \text{ cm}^{-3}$ 

**Applications** 

• Ideal **model system** to study the physics of ultrahigh intensity laser-plasma interaction • Optical elements to manipulate extreme laser intensities

• New sources of ultrashort pulses of light or particles at high energies

## Attosecond pulses from plasma mirrors

The non-linear PM response produces high-order harmonics, associated to trains of attosecond pulses in the time domain



Plaja et al, J. Opt. Soc. Am. B 15(1998)

## High-energy particles from plasma mirrors

#### Beams of relativistic electrons and high-energy ions are also produced



#### Thévenet et al, Nature Phys. 12, 355 (2016)

## A little bit of history: first HHG experiments



# **Observation of the** broad high-order harmonic radiation in gas targets

## A little bit of history: first HHG experiments

#### Laser HHG started in dense plasmas (NOT plasma mirrors?)

Los Alamos National Lab, early 80's

HHG from solid targets with intense far-infrared, nanosecond <u>CO<sub>2</sub> lasers</u> (λ=10 μm)

 $I \lambda^2 \rightarrow 10^{18} \text{ W/cm}^2 \, \mu\text{m}^2$ 

#### Gemini laser (power amplifier exit end)



Burnett et al. Appl. Phys. Lett. 31 (3): 172–174 (1977) R.L. Carman et al, Phys. Rev. Lett. **46** (1981); Phys. Rev. A **24** (1981)



#### Some promising early numerical and experimental results

Gordienko et al, Phys. Rev. Lett. 93 (2004) Dromey et al, Nature phys. 2 (2006) & Phys. Rev. Lett. 99 (2007) **PIC** simulations Experiment Photon Energy, eV  $10^{1}$  $I \sim \omega^{-5/2}$ 1770 2360 2950 \_=10 3530 Intensity, a.u. 1010 Intensity/arb. units Normalised at 1200<sup>th</sup> order  $10^{8}$ *n<sub>B0</sub>*≈ 2600 106 *n<sub>R0</sub>≈*\_3000 104  $10^{2}$ 1000 10 100  $\omega_n / \omega_0$ a)  $(1.5\pm.0.3)\times10^{20}$ Wcm<sup>-2</sup> p=2.8 -b) (2.5±.0.5)×10<sup>20</sup>Wcm<sup>-2</sup> Tsakiris et al, New J. Phys. 8 (2006) Prel=2.55 (+0.25, -0.15) **10**<sup>-2</sup> **PIC** simulations 2500 1500 2000 3000 Harmonic order, n 10°  $l = \lambda/4$ For a review until  $\approx 2008$ See Teubner & Gibbon, Rev. Mod. Phys. 81 (2009) effici 10<sup>-5</sup> Zt 10<sup>-6</sup> (80-200eV) **VULCAN** 10 Cu @RAL (UK) (400-1000-eV) 10 ≈1 PW - 600 fs 10 100  $a_{L}$ 

## Outline

#### 1- What tools?

- Particle-In-Cell (PIC) codes
- Experimental tools
  - $\rightarrow$  Plasma mirrors for contrast improvement

#### 2- HHG: basic physical mechanisms

- Relativistic oscillating mirror (ROM)
- Coherent wake emission (CWE)

#### 3- Control (and metrology) of harmonic emission

- Controlling the interface steepness
- Transient plasma gratings (& plasma holograms)
- Attosecond lighthouses

## Particle-in-Cell codes, a major tool for UHI physics





#### 'UHI100' @ CEA-IRAMIS

<u>P = 100 TW</u> - E=2.5 J - τ=25 fs − 10 Hz Final beam aperture ≈80 mm,  $w_0 \approx 4 \mu m$  $I\lambda^2 \approx 5.10^{19} W cm^{-2} \mu m^2$ 



## Beam conditionning



## The issue of the temporal contrast of ultrashort lasers



Thin foil probed 1 ns <u>before</u> The main pulse, Already destroyed !!!



## Optical switching using plasma mirrors



H.C. Kapteyn et al, Opt. Lett. 16 490 (1991)



## After the double plasma mirror...

#### **Overall transmission of DPM : 50 %** Duration and wavefront unaltered



## Plasma mirrors in action



## Plasma mirror after some shots





## HHG: basic physical mechanisms

## **Relativistic Oscillating Mirror**



Lichters et al, Phys. Plasmas 3 (1996)

## ROM observed in simulations

#### Particle-in-Cell simulation: I=1.5 10<sup>19</sup> W/cm<sup>2</sup> - L= $\lambda/8$



Harmonic generation with a 1 TW-50 fs laser system (LUCA)



F.Quéré et al, Phys. Rev. Lett. 96, 125004 (2006)

## We tried an experiment that shouldn't have worked

yet it did work, and from it we learned a lot of physics !

## Lesson to remember:

In experiments,

you should not always look for the effects you expect,

but <u>also</u> -sometimes- for things you absolutely do not expect

## **Back to PIC simulations**



## Coherent Wake Emission (CWE)



## **Experimental evidence: CWE & ROM**

## Similarity with short and long trajectories signals in gas HHG



Thaury et al, Nature Physics **3**, 424 (2007)

## Summary: mechanisms and harmonic properties

#### **Relativistic Oscillating Mirror**

- Doppler effect
- Harmonic cut-off depends on laser intensity
- *Requires highest possible intensities (>10<sup>18</sup> W/cm<sup>2</sup>.µm<sup>2</sup>)*
- Attosecond (zepto?) pulses close to their Fourier limit

#### **Coherent Wake Emission**

- Linear mode conversion from plasma oscillations triggered by electron bunches
- Harmonic spectral cut-off = maximum plasma frequency  $\alpha$  (plasma density)<sup>1/2</sup>
- Only requires moderate intensities, >10<sup>16</sup> W/cm<sup>2</sup>.µm<sup>2</sup>
- Slightly chirped attosecond pulses











Control and metrology of harmonic emission



## Controlling and measuring the interface steepness







## CWE to ROM transition for varying interface steepness

## **I=10<sup>18</sup> W/cm<sup>2</sup>**

**Experimental results** 



Laser UHI 100 (CEA) 25 fs 100 TW

Kahaly et al PRL **110** (2013)

## Transition to chaotic dynamics



Chopineau et al, Phys. Rev. X 9, 011050 – (2019) Blaclard et al, Phys. Rev. E 107, 034205 (2023)

## Transient plasma gratings: key idea



## Transient plasma gratings: key idea





#### Plasma gratings

Monchocé et al, Phys. Rev. Lett. 112, 145008 (2014)

## Ptychographic measurement of the harmonic source spatial profile (amplitude and phase)

Leblanc et al, Nature Physics **12**, 301–305 (2016) Leblanc et al, Phys. Rev. Lett. **119**, 155001 (2017)

#### Plasma holograms

Leblanc et al, Nature Physics 13, 440–443 (2017)



Intense femtosecond pulse

Train of attosecond pulses

Plasma mirror

## Spatio-temporal control: the attosecond lighthouse effect



## Focusing a 'normal' pulse



## Focusing a pulse with pulse front tilt



## **Experimental demonstration**



Footprint of the XUV « harmonic » beam in the far field as a function of the laser pulse **CEP** 

Wheeler et al, *Nature Photonics* **6**, 828-832 (2012) Kim et al, Nature Photonics 7, 651–656 (2013)



## Advanced metrology: attosecond temporal measurements



## Many were topics not covered here



- ✓ Different theoretical models of relativistic harmonic generation, and associated controversy
- ✓ Spatial and spectral phase properties of harmonics and associated models
- ✓ Approaches for spatial and temporal metrology, e.g. ptychography
- $\checkmark$  Optimization and control of harmonic emission ( $\omega/2\omega$ , CEP, vortex beams....)
- ✓ Temporal gating techniques for the generation of isolated attosecond pulses
- Electron acceleration: using plasma mirrors as injectors for Vacuum Laser Acceleration or laser wakefield acceleration
- ✓ Transition to chaotic dynamics when the plasma interface gets smoother

#### **Conclusion & perspectives**

## Considerable progress in the last $\approx$ 15 years

- Good understanding of the harmonic generation mechanisms
- Major advance in control and metrology of harmonics/atto pulses

→ Rich physics, insight into ultrahigh intensity interactions

→ Future attosecond sources complementary to HHG in gases?

→ Developments of more compact ultraintense laser sources, higher rep rates, new target technologies

[1 kHz, 1.5-cycle, 780 nm, 1 TW ] @

[1 kHz, 3-cycle, 900 nm, 5-15 TW ] @





Y. H. Kim et al., Nature Comms. 14, 2328 (2023)

## SHHG beamline @ ELI-ALPS

#### SourceLAB | Laser Plasma Technologies



## SHHG beamline @ ELI-ALPS



## Fundamental physics with PW lasers?

#### What questions in fundamental physics can be addressed with high-power lasers?



- Intensities  $> 10^{25}$  W/cm<sup>2</sup>-10<sup>29</sup> W/cm<sup>2</sup> are needed
- The present record in laser intensity is 'only' ≈10<sup>23</sup> W/cm<sup>2</sup> Yoon et al, Optica **8**, 630–635 (2021)

## Potential solution: reflection off curved relativistic mirror

## → The Curved Relativistic Mirror (CRM) concept



(i) Intensification by temporal compression
Landecker, 86, 852 Phys. Rev. (1952)
(ii) Intensification by spatial focusing to a tighter spot (λ << λ<sub>L</sub>)
Bulanov et al, PRL 91, 095001 (2003)

But how to actually implement this in the lab? ⇒ This might be achieved with plasma mirrors

#### Validation by 3D PIC simulations: case of a 3 PW laser

**3D pseudo-spectral PIC simulation with WARP-PICSAR (≈20.10<sup>6</sup> CPU hours)** → INCITE program - MIRA supercomputer @ Argonne lab



Compressed Atto pulse: 5.5J, 100as, 350nm  $\rightarrow$  I=10<sup>25</sup>W/cm<sup>2</sup> Only 30 harmonic orders contribute to the intensity gain !

## Relativistic plasma mirrors : a feasible implementation of a CRM



#### What are the maximum intensities achievable with this scheme?



#### Achievable intensities with curved relativistic plasma mirrors



## Contributors



#### CEA

Cédric Thaury Henri Vincenti Sylvain Monchocé Adrien Leblanc Ludovic Chopineau Guillaume Bouchard Guillaume Blaclard Subhendu Kahaly Adrien Denoeud Philippe Martin

#### Collaborators

<u>LOA</u> R. Lopez-Martens J. Faure M. Thévenet J. Wheeler A. Borot A.Malvache

> <u>LULI</u> J-P. Geindre

> > <u>DPTA</u> L. Videau P. Combis

#### *Laser operation (CEA) F. Réau C. Pothier*

D. Garzella P. d'Oliveira



