



ELISSS2023

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Dolní Břežany, Czech Republic

Electromagnetic Particle-In-Cell simulations

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IMPULSE

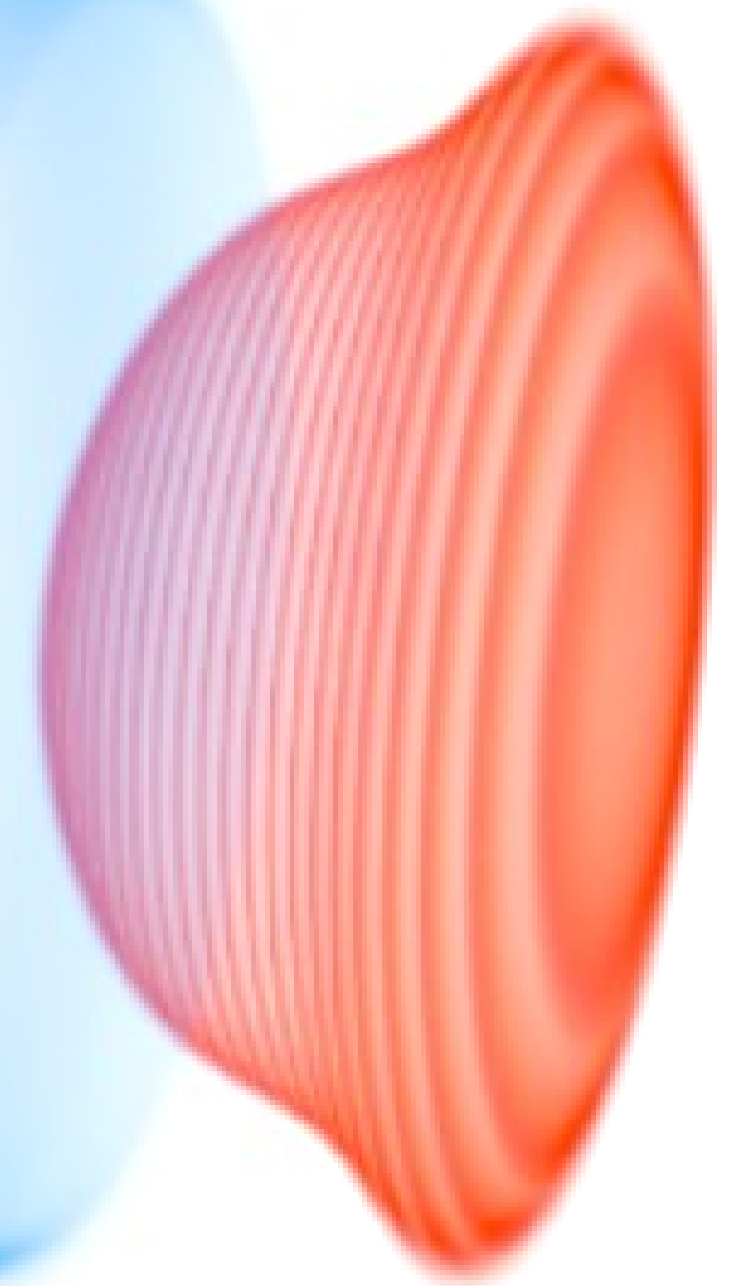


IMPULSE is funded by the European Union's Horizon 2020 programme under grant agreement No. 871161

Electromagnetic Particle-In-Cell simulations

Kinetic description of plasma
(not hydrodynamic)

Electromagnetic effects
(not only electrostatic)



An *Electromagnetic* PIC code:

- For which purpose?
- How does it work?
- Is it stable?
- What physics can it handle?
- On which computer can it work?

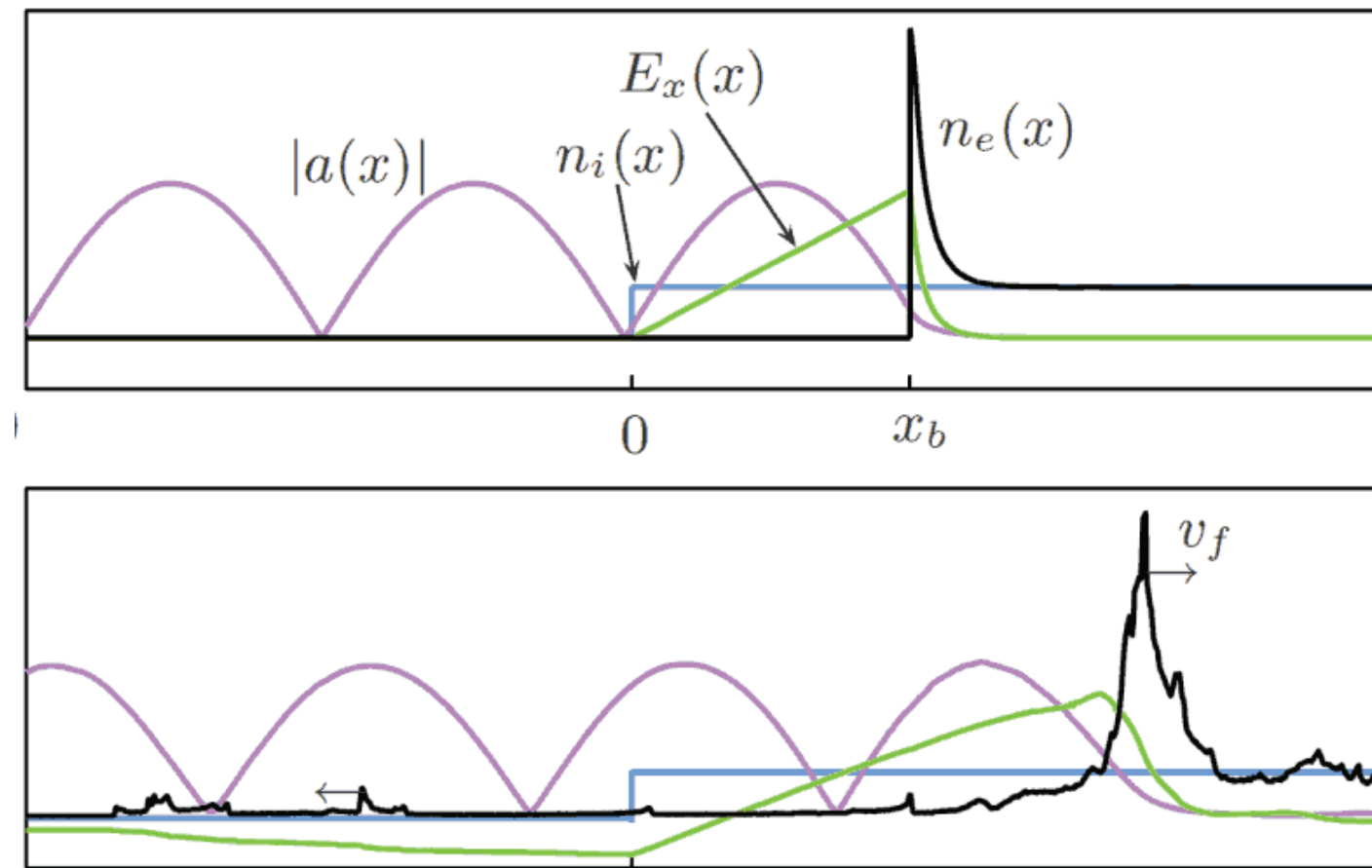
An *Electromagnetic* PIC code:

- **For which purpose?**
- How does it work?
- Is it stable?
- What physics can it handle?
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PIC codes are an excellent support to theoretical modeling (even in 1D)

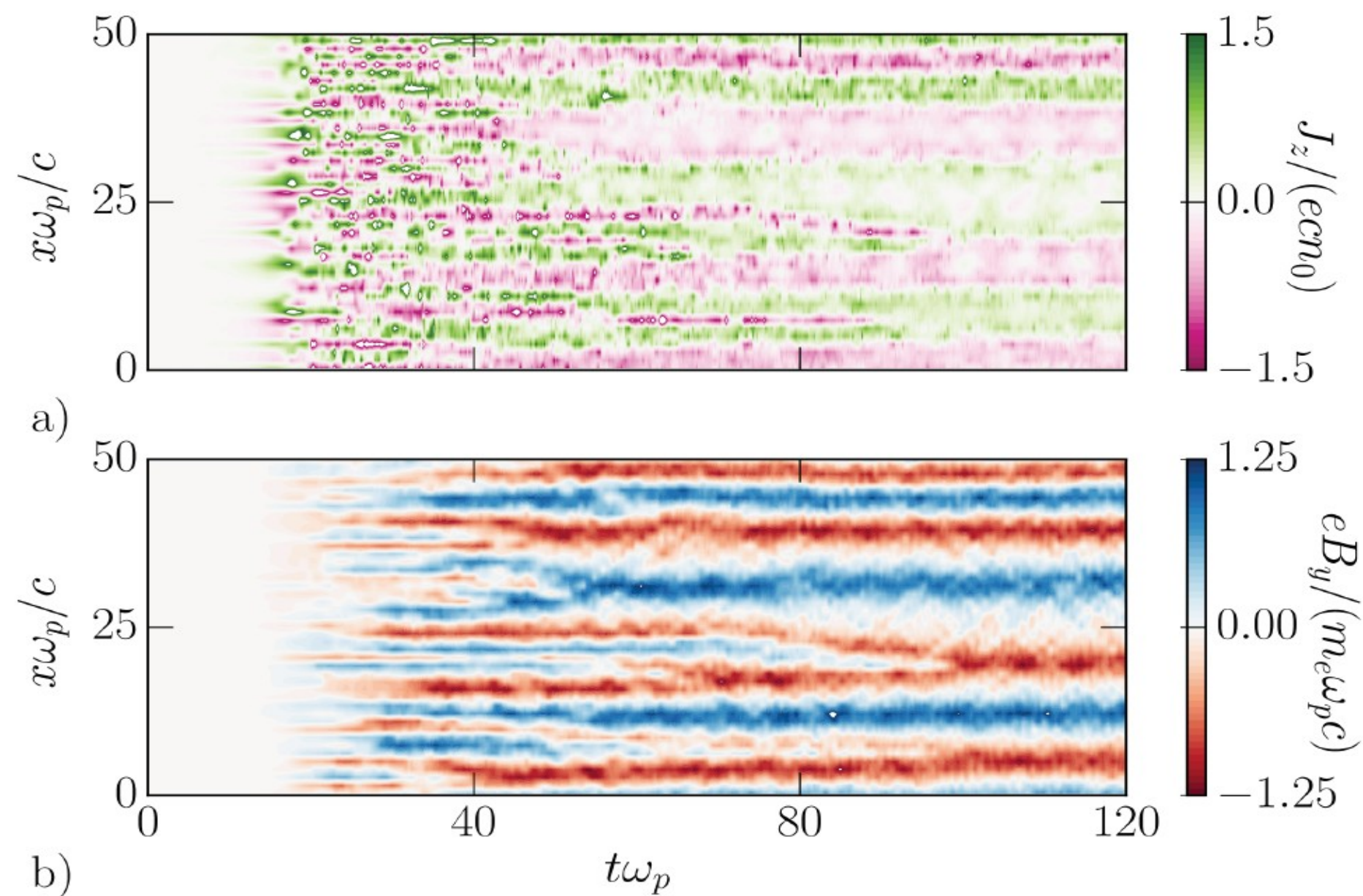
Relativistically-induced transparency

Siminos et al., PRE 86, 056404 (2012)



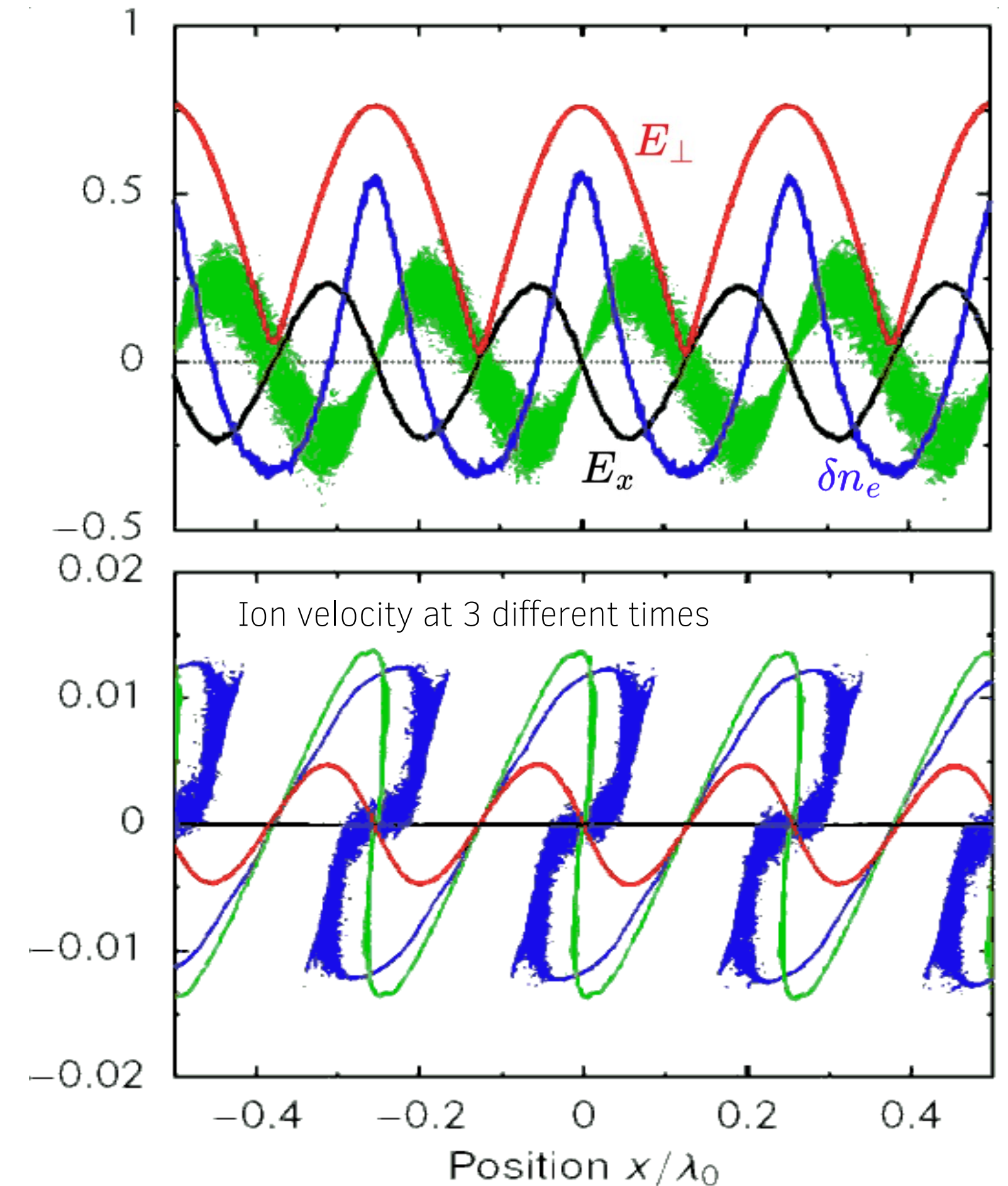
Weibel instability in the presence of an external magnetic field

Grassi et al., PRE 95, 023203 (2017)



Standing Whistler Waves

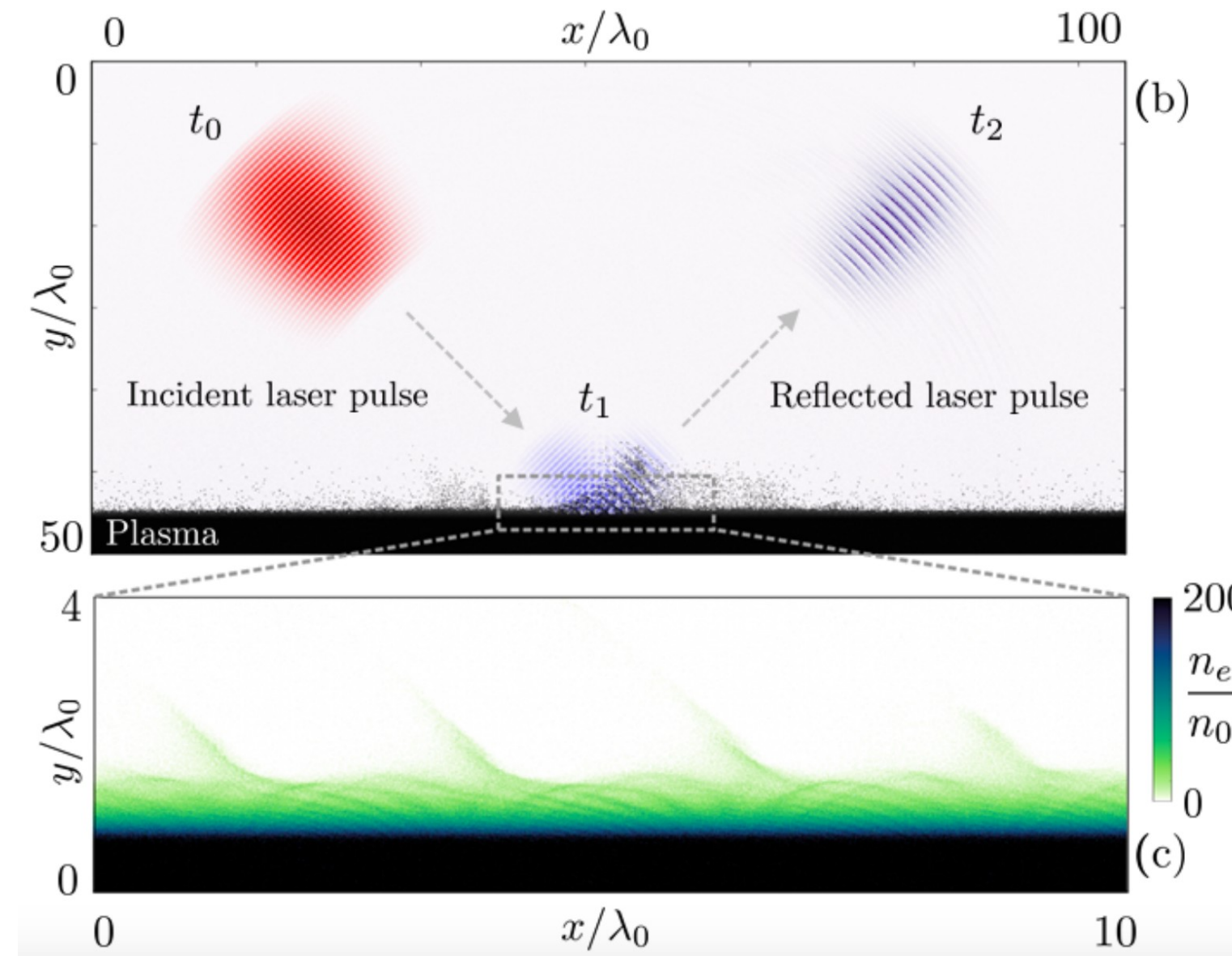
Sano et al., PRE 100, 053205 (2019)



2D or 3D simulations help design or understand experiments

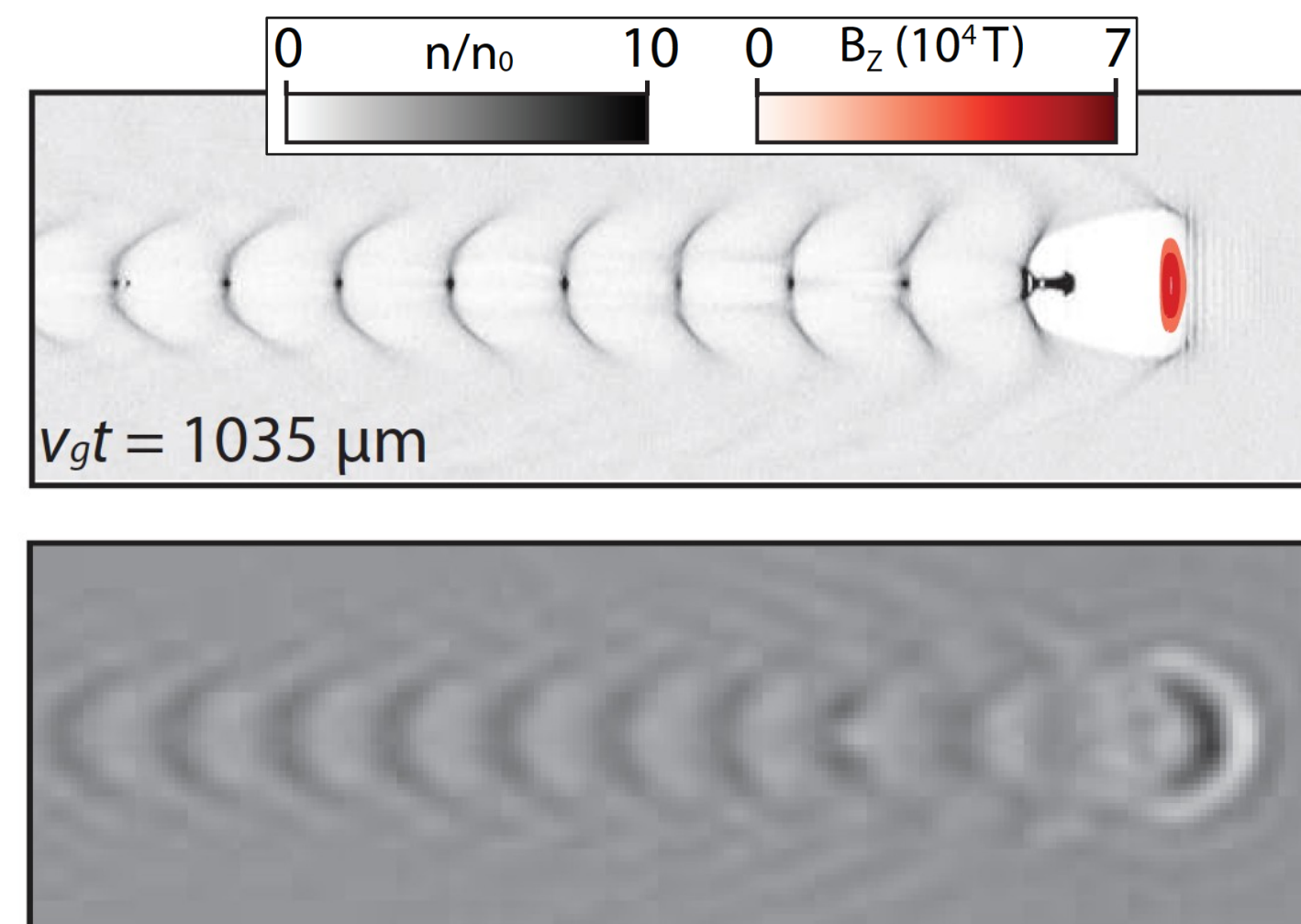
High-harmonic generation

G. Bouchard, F. Quéré,
CEA/IRAMIS



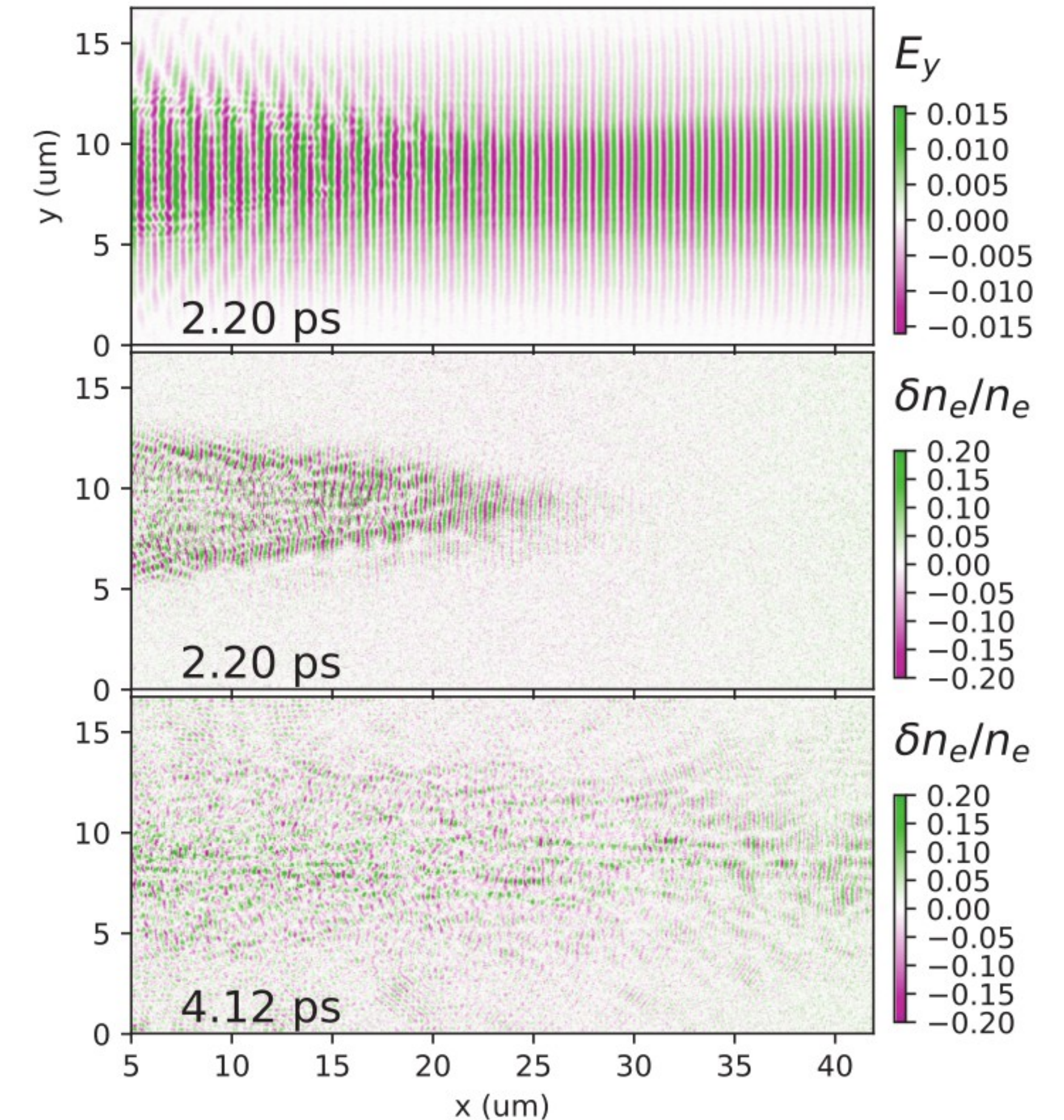
Laser wakefield acceleration

Sävert *et al.*,
PRL **115**, 055002 (2015)



Plasma instabilities induced by laser

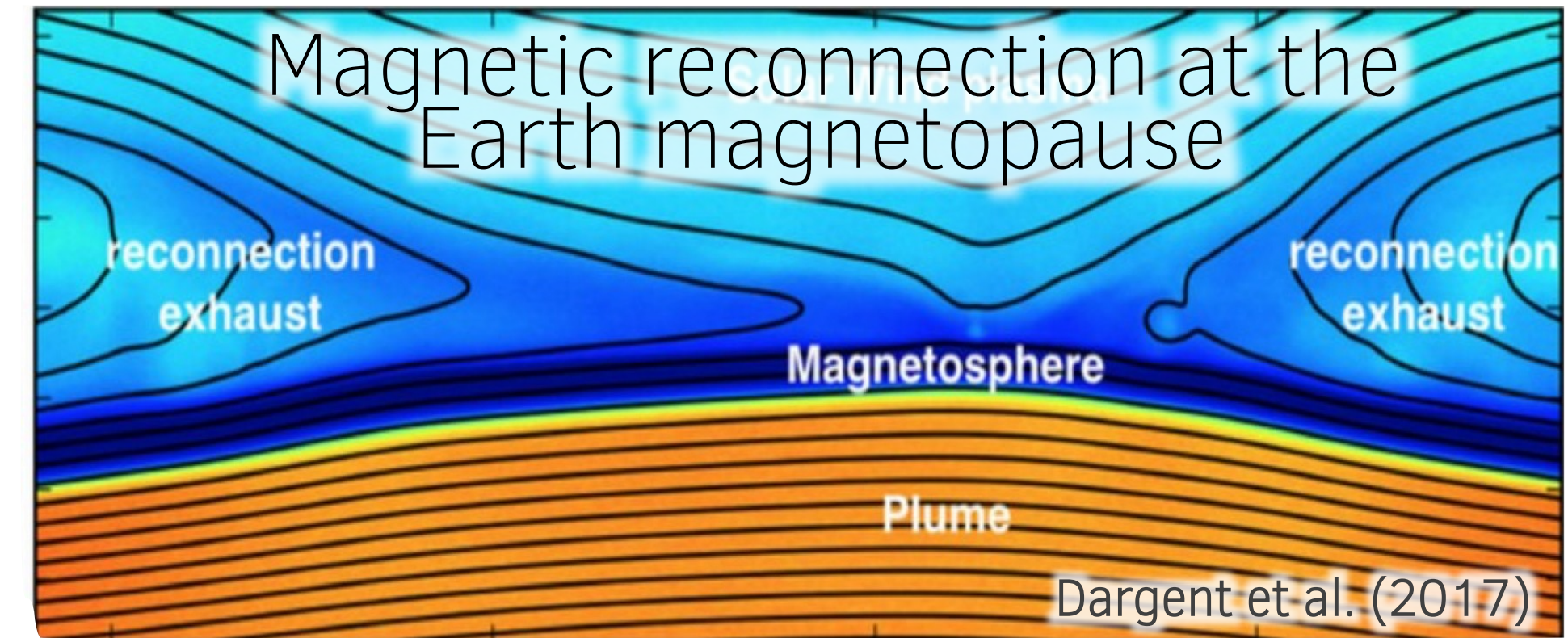
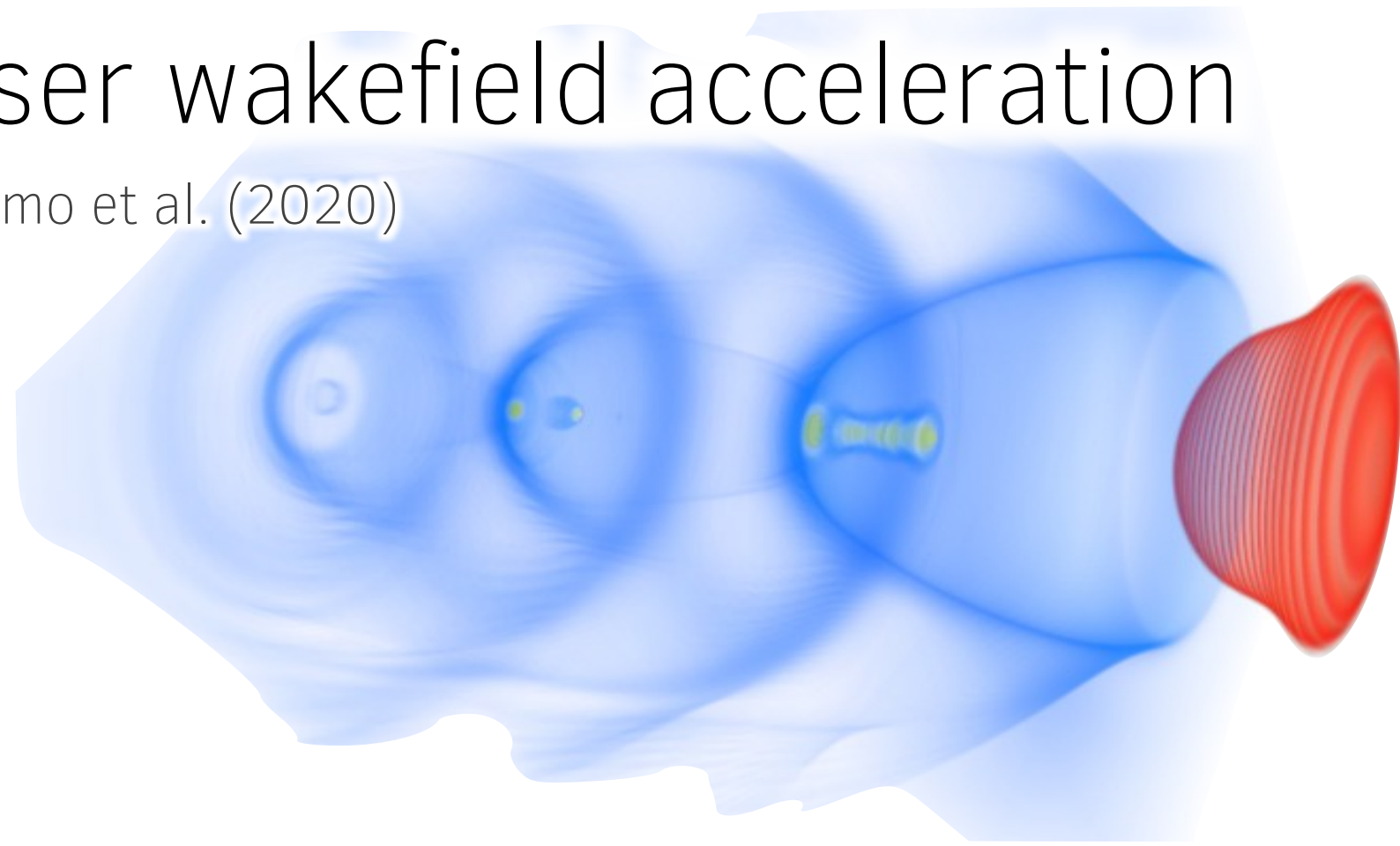
Perez *et al.* PoP **28**, 043102 (2021)



Overall, electromagnetic PIC codes are central to a wide range of plasma-physics-related studies

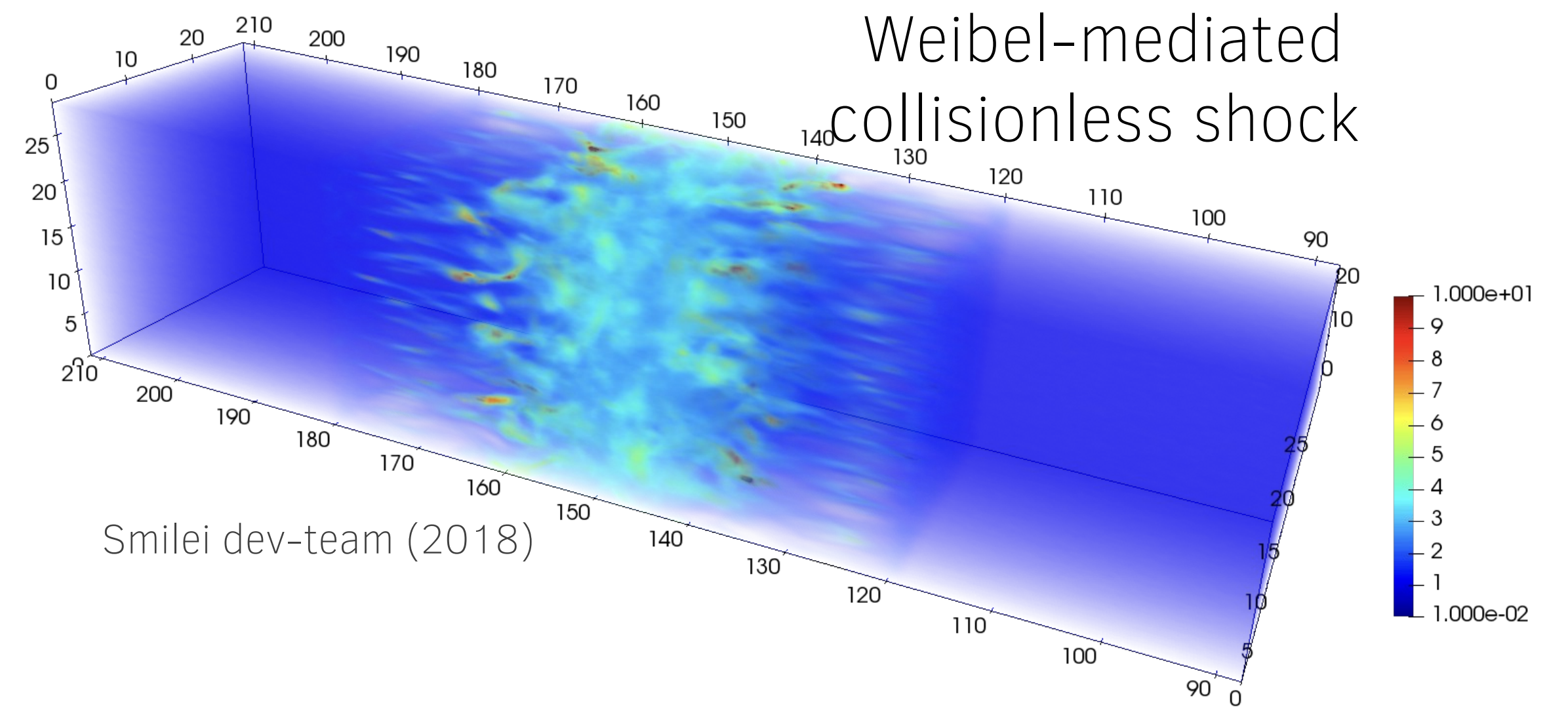
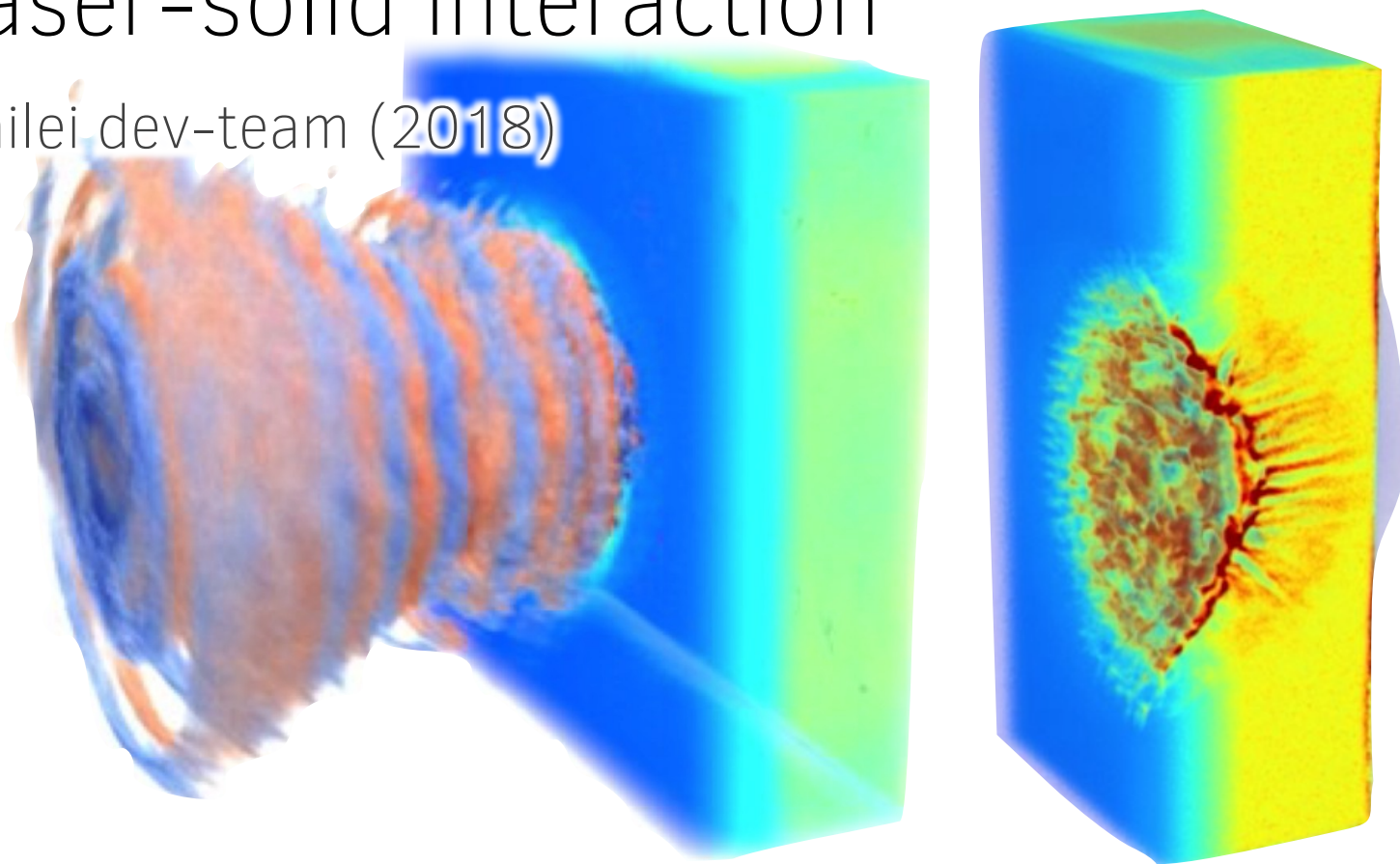
Laser wakefield acceleration

Massimo et al. (2020)



Laser-solid interaction

Smilei dev-team (2018)



+ plasma propulsion, cosmology, ...

An *Electromagnetic* PIC code:

- For which purpose?
- **How does it work?**
- Is it stable?
- What physics can it handle?
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Quick reminder: special relativity

Momentum

$$\mathbf{p} = \gamma m \mathbf{v}$$

Newton's
2nd law

$$\frac{d\mathbf{p}}{dt} = \mathbf{F}$$

Lorentz
factor

$$\gamma = \frac{1}{\sqrt{1 - (\mathbf{v}/c)^2}} = \sqrt{1 + (\mathbf{p}/(mc))^2}$$

A kinetic description of plasma

$f_s(t, \mathbf{x}, \mathbf{p})$ Distribution function

species

density

$$\rho(t, \mathbf{x}) = \int d^3\mathbf{p} f_s(t, \mathbf{x}, \mathbf{p})$$

current density

$$\mathbf{J}(t, \mathbf{x}) = q_s \int d^3\mathbf{p} \mathbf{v} f_s(t, \mathbf{x}, \mathbf{p})$$

$$\partial_t f_s + \mathbf{v} \cdot \nabla f_s + \mathbf{F}_L \cdot \nabla_{\mathbf{p}} f_s = 0$$

Vlasov's
equation

Average distribution

Average Lorentz force

Add Fields and Maxwell's equations

$\mathbf{E}(t, \mathbf{x})$ Electric field

$\mathbf{B}(t, \mathbf{x})$ Magnetic field

→ Lorentz force $\mathbf{F}_L = q_s (\mathbf{E} + \mathbf{v} \times \mathbf{B})$

$$\nabla \cdot \mathbf{E} = \frac{\rho}{\epsilon_0}$$

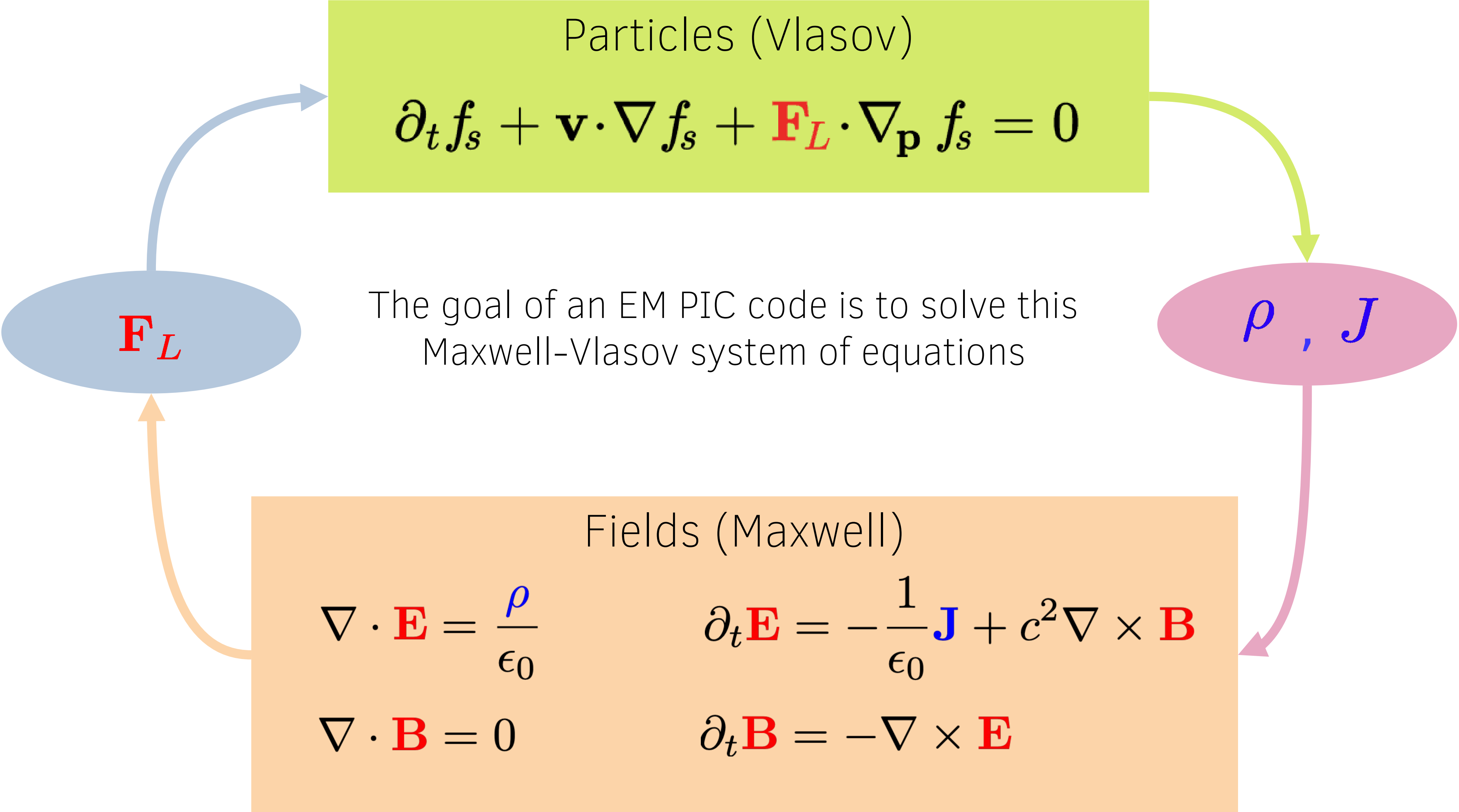
$$\partial_t \mathbf{E} = -\frac{1}{\epsilon_0} \mathbf{J} + c^2 \nabla \times \mathbf{B}$$

$$\nabla \cdot \mathbf{B} = 0$$

$$\partial_t \mathbf{B} = -\nabla \times \mathbf{E}$$

Maxwell's
equations

Combine Vlasov + Maxwell



How to solve Vlasov + Maxwell numerically?

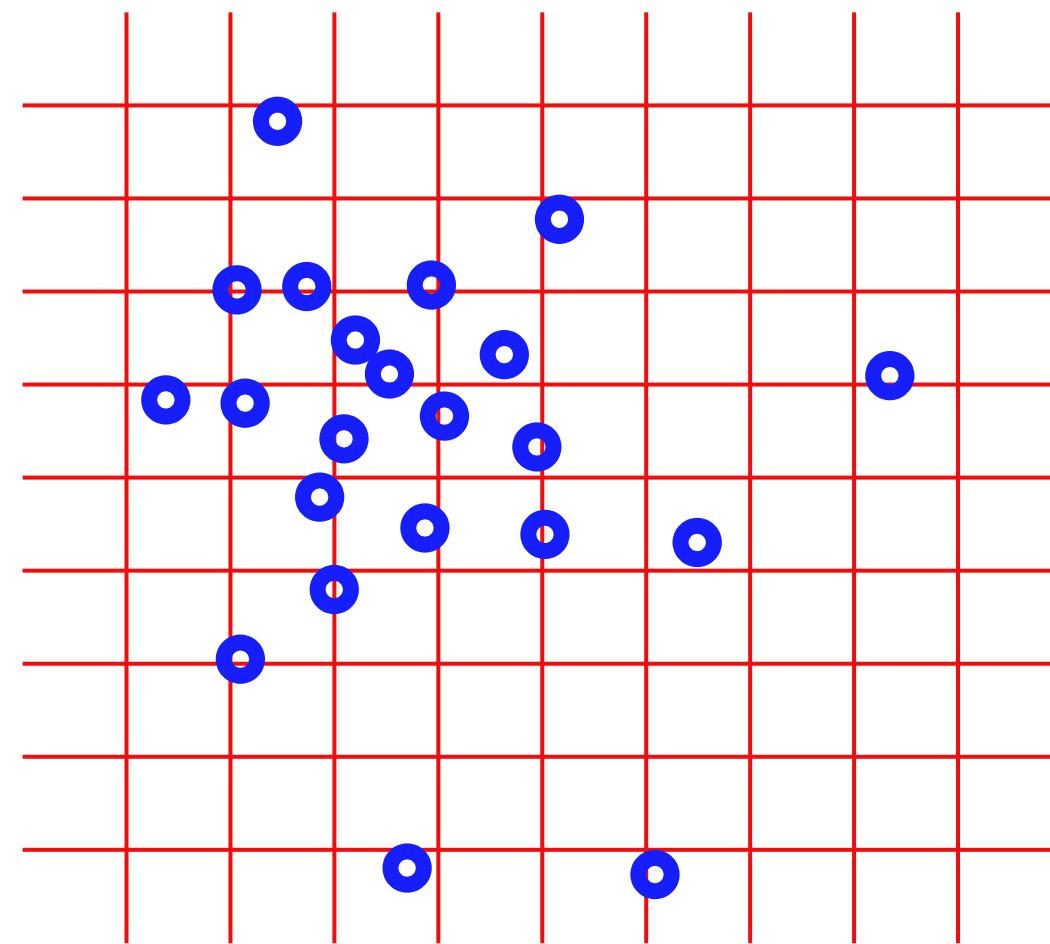
Particle-in-Cell

Vlasov

quasi-particles

Maxwell

discrete grid for fields



Distribution function = sum of quasi-particles

$$\partial_t f_s + \mathbf{v} \cdot \nabla f_s + \mathbf{F}_L \cdot \nabla_{\mathbf{p}} f_s = 0$$

Vlasov = partial differential equations in a 6D space.

Direct integration (*see Vlasov codes*) has tremendous computational cost!

In a PIC code, the distribution function is approximated as a sum over quasi-particles

$$f_s(t, \mathbf{x}, \mathbf{p}) = \sum_{p=1}^N w_p S(\mathbf{x} - \mathbf{x}_p(t)) \delta(\mathbf{p} - \mathbf{p}_p(t))$$

Statistical weight

Shape function

Now, inject this distribution in Vlasov's equation ...

PIC distribution + Vlasov equation = quasi-particle motion

Integrate over momentum $\rightarrow \partial_t \mathbf{x}_p = \mathbf{v}_p$

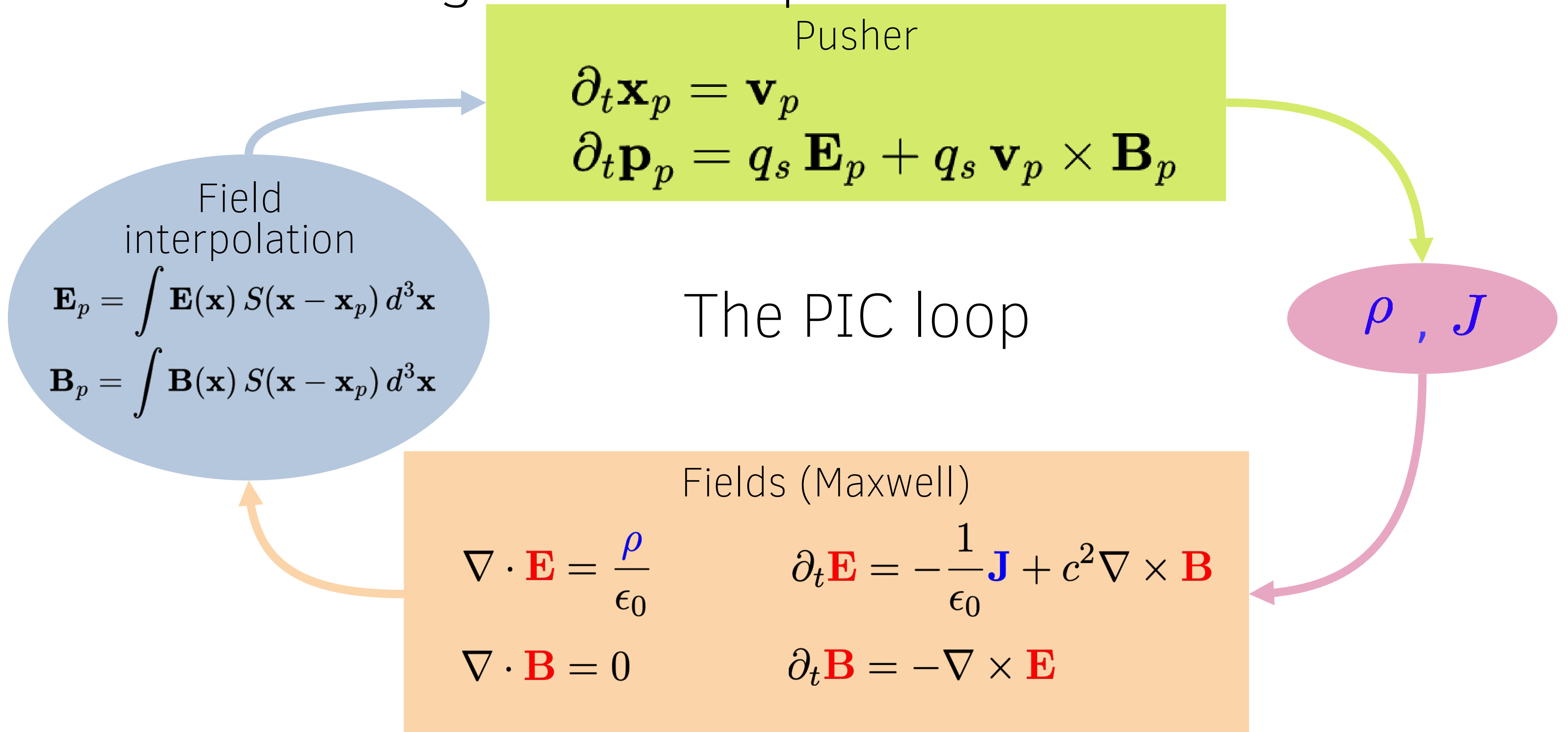
Multiply by \mathbf{p} then integrate over momentum and space $\rightarrow \partial_t \mathbf{p}_p = q_s \mathbf{E}_p + q_s \mathbf{v}_p \times \mathbf{B}_p$

$$\text{where } \begin{cases} \mathbf{E}_p = \int \mathbf{E}(\mathbf{x}) S(\mathbf{x} - \mathbf{x}_p) d^3 \mathbf{x} \\ \mathbf{B}_p = \int \mathbf{B}(\mathbf{x}) S(\mathbf{x} - \mathbf{x}_p) d^3 \mathbf{x} \end{cases}$$

Fields interpolated at the location of quasi-particles.
Note that the fields \mathbf{E} and \mathbf{B} exist only at grid points.

The movement of quasi-particles is essentially that of real particles

Replacing Vlasov by the quasi-particle motion & introducing field interpolation



Maxwell's equations can be simplified

Assume Maxwell-Ampere, Maxwell-Faraday and conservation of charge.

$$\partial_t \mathbf{E} = -\frac{1}{\epsilon_0} \mathbf{J} + c^2 \nabla \times \mathbf{B} \quad \partial_t \mathbf{B} = -\nabla \times \mathbf{E} \quad \partial_t \rho + \nabla \cdot \mathbf{J} = 0$$

Take the divergence

$$\partial_t \nabla \cdot \mathbf{E} + \nabla \cdot \mathbf{J} = 0$$

$$\partial_t (\nabla \cdot \mathbf{E} - \rho) = 0$$

$$\partial_t (\nabla \cdot \mathbf{B}) = 0$$

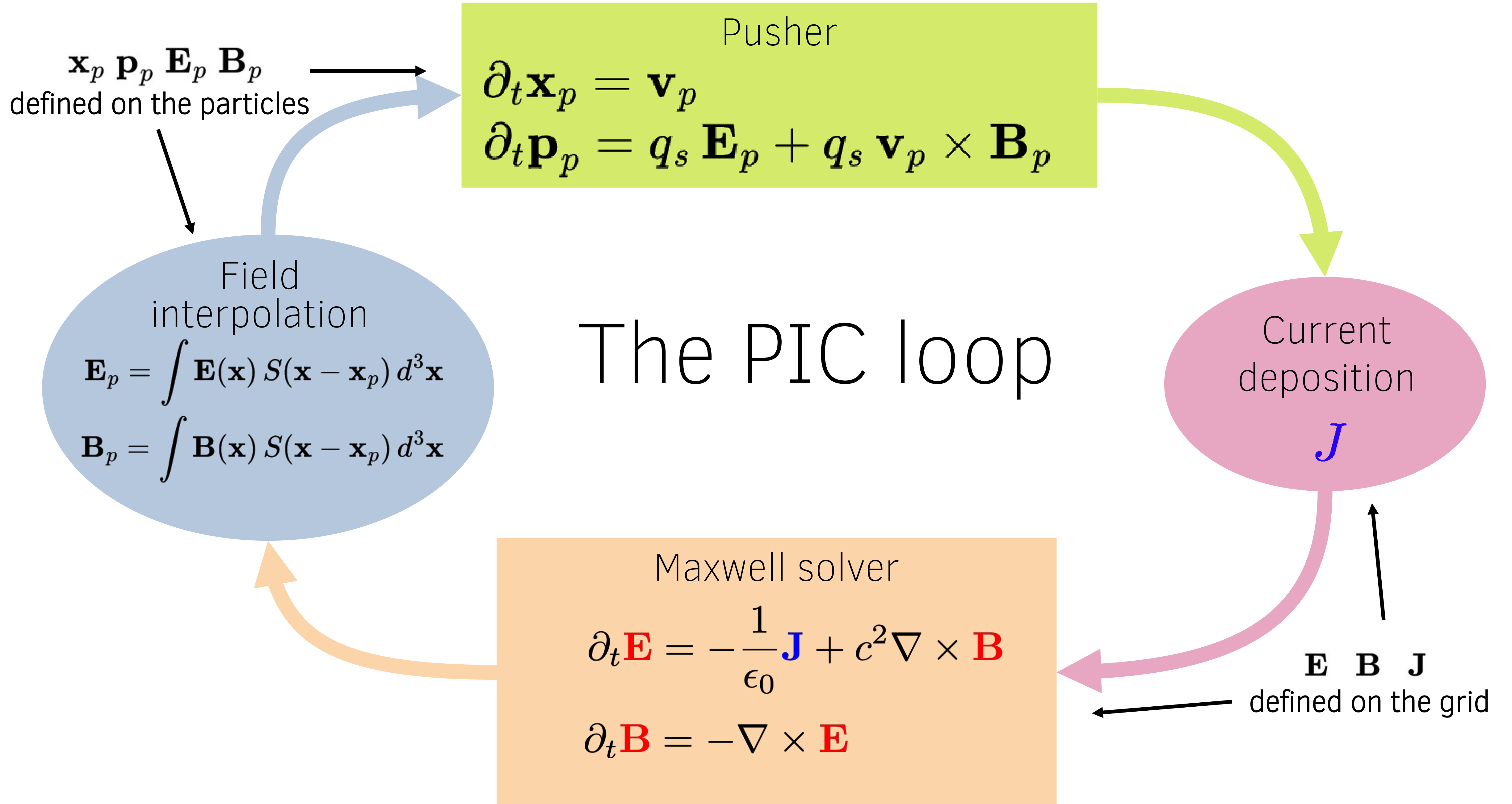
└───────────> Maxwell-Gauss conserved

└───────────> Maxwell-Poisson conserved

If Maxwell-Gauss and Maxwell-Poisson are satisfied initially,

it is not necessary to solve them.

(Then it is also unnecessary to calculate the density)



In practice, many other aspects must be considered

- System of units
- Shape function \mathcal{S}
- Geometry of the grid
- Ensure conservation of charge
- Boundary conditions
- Initial conditions
- ...

The natural units system

Velocity	c
Charge	e
Mass	m_e
Momentum	$m_e c$
Energy, Temperature	$m_e c^2$
Time	ω_r^{-1}
Length	c/ω_r
Number density	$n_r = \epsilon_0 m_e \omega_r^2 / e^2$
Current density	$e c n_r$
Pressure	$m_e c^2 n_r$
Electric field	$m_e c \omega_r / e$
Magnetic field	$m_e \omega_r / e$
Poynting flux	$m_e c^3 n_r / 2$

ω_r is the reference angular frequency. It does not need to be defined in the code. Results of the simulation can be scaled *a posteriori* by changing the value of ω_r .

Fields (Maxwell)

$$\nabla \cdot \mathbf{E} = \rho \quad \partial_t \mathbf{E} = -\mathbf{J} + \nabla \times \mathbf{B}$$

$$\nabla \cdot \mathbf{B} = 0 \quad \partial_t \mathbf{B} = -\nabla \times \mathbf{E}$$

Maxwell's equations are simplified when all quantities are normalized with the natural units.

Choose the shape function \mathcal{S}

$$\hat{s}^{(0)}(x) = \Delta x \delta(x),$$

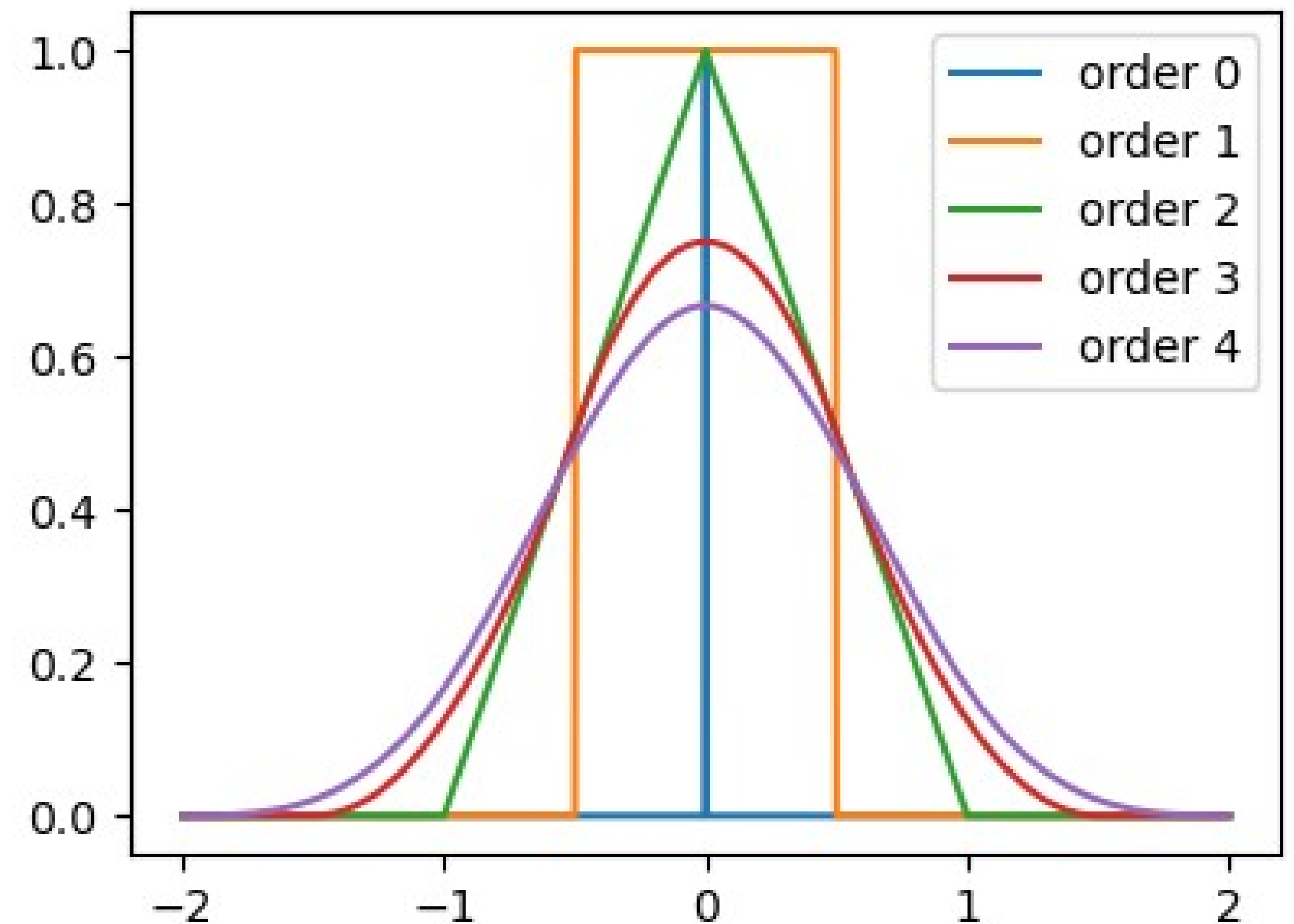
$$\hat{s}^{(1)}(x) = \begin{cases} 1 & \text{if } |x| \leq \frac{1}{2} \Delta x, \\ 0 & \text{otherwise,} \end{cases}$$

$$\hat{s}^{(2)}(x) = \begin{cases} \left(1 - \left|\frac{x}{\Delta x}\right|\right) & \text{if } |x| \leq \Delta x, \\ 0 & \text{otherwise,} \end{cases}$$

$$\hat{s}^{(3)}(x) = \begin{cases} \frac{3}{4} \left[1 - \frac{4}{3} \left(\frac{x}{\Delta x}\right)^2\right] & \text{if } |x| \leq \frac{1}{2} \Delta x, \\ \frac{9}{8} \left(1 - \frac{2}{3} \left|\frac{x}{\Delta x}\right|\right)^2 & \text{if } \frac{1}{2} \Delta x < |x| \leq \frac{3}{2} \Delta x, \\ 0 & \text{otherwise,} \end{cases}$$

$$\hat{s}^{(4)}(x) = \begin{cases} \frac{2}{3} \left[1 - \frac{3}{2} \left(\frac{x}{\Delta x}\right)^2 + \frac{3}{4} \left|\frac{x}{\Delta x}\right|^3\right] & \text{if } |x| \leq \Delta x, \\ \frac{4}{3} \left(1 - \frac{1}{2} \left|\frac{x}{\Delta x}\right|\right)^3 & \text{if } \Delta x < |x| \leq 2 \Delta x, \\ 0 & \text{otherwise.} \end{cases}$$

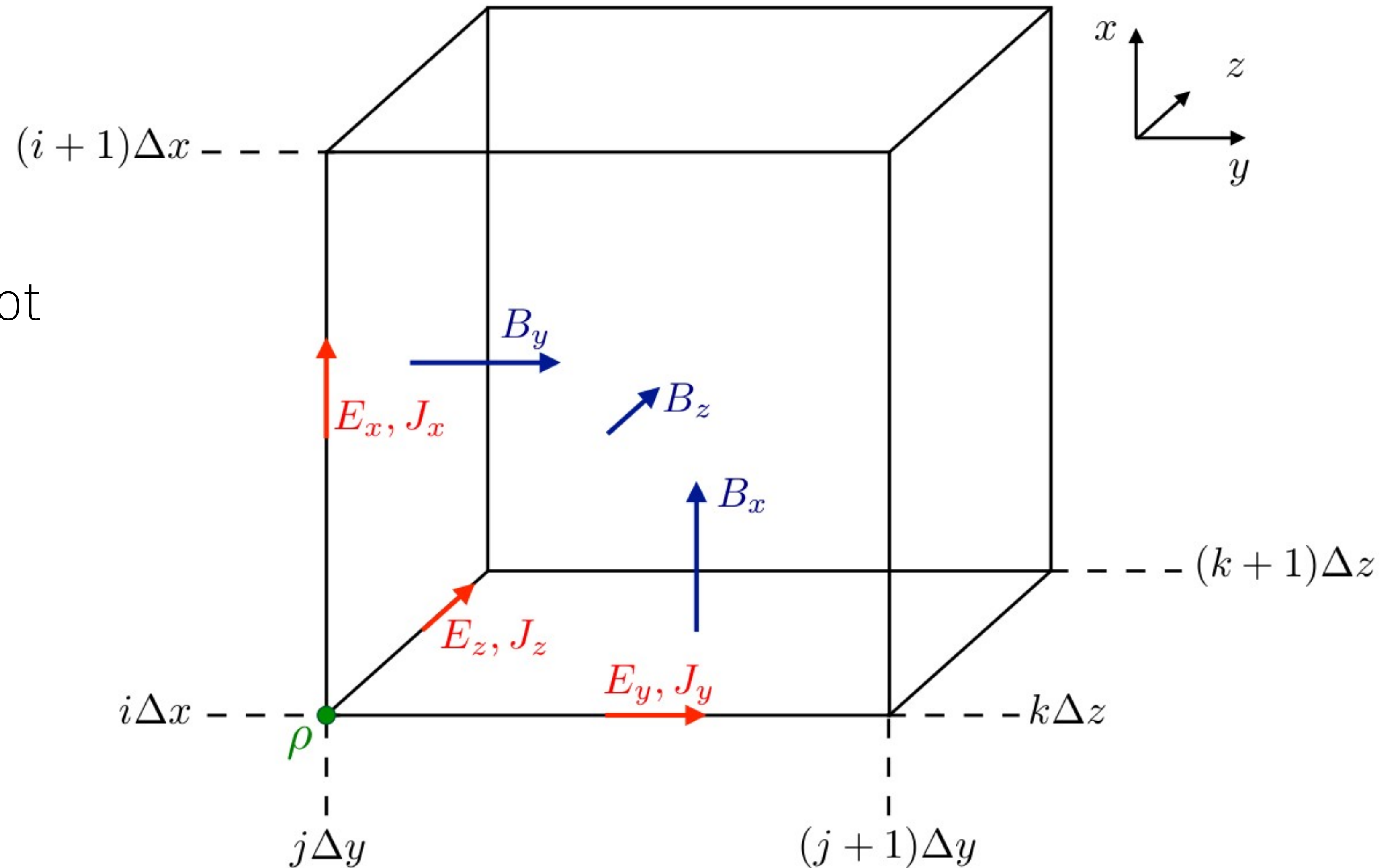
The shape function may be of different orders that increase the smoothing of field interpolation. A higher order reduces noise, but longer to compute.



Finite-difference time-domain (FDTD) (or Yee's grid)

A popular method to solve Maxwell

Fields and currents are not defined on the nodes.



The PIC algorithms may also take many different forms

- Cylindrical or spherical geometry
- Various models for pusher (ex: implicit)
- Various models for solver (ex: spectral, envelope)
- ...

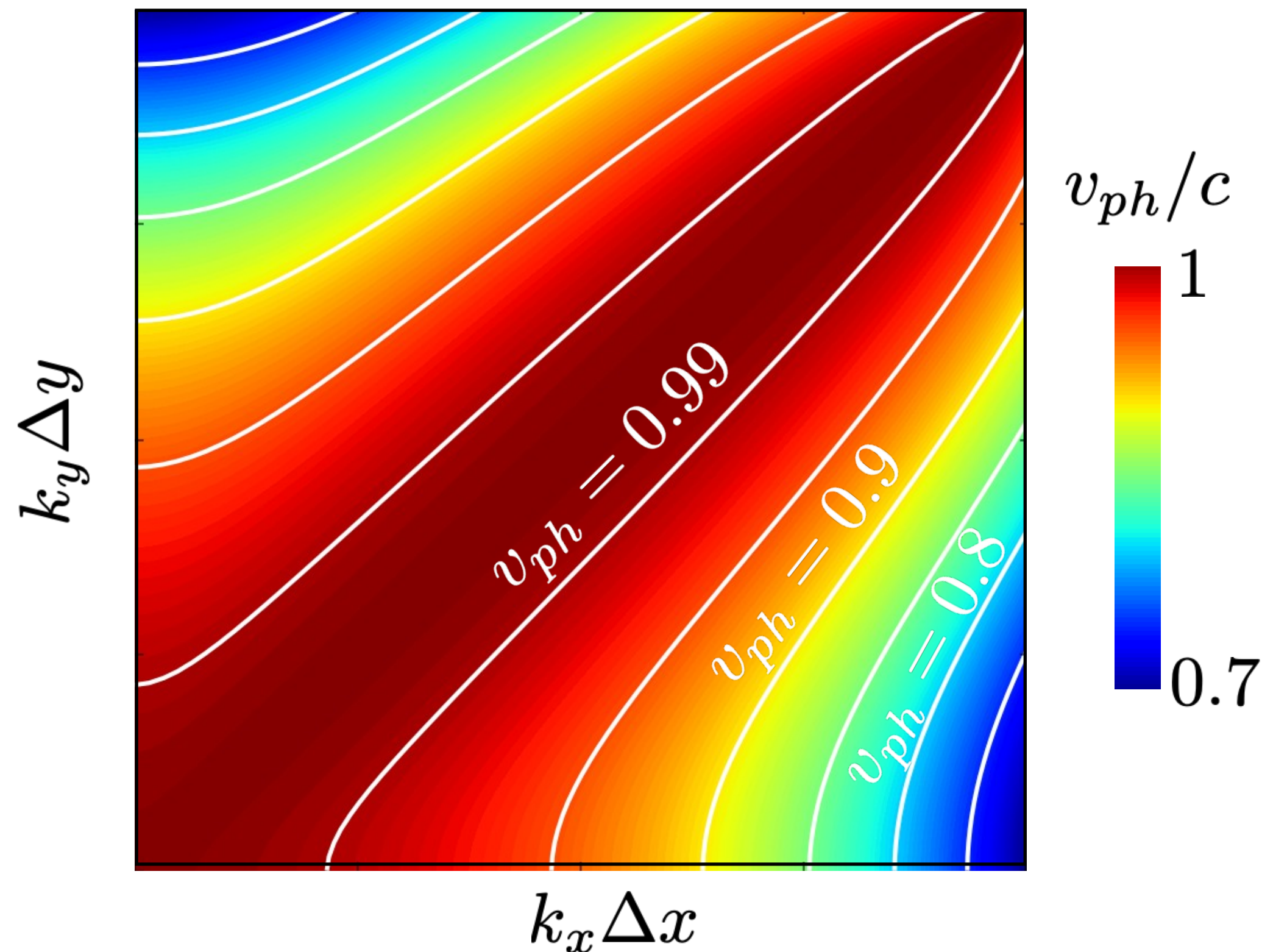
An *Electromagnetic* PIC code:

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Numerical analysis of the FDTD solvers shows some of their limitations

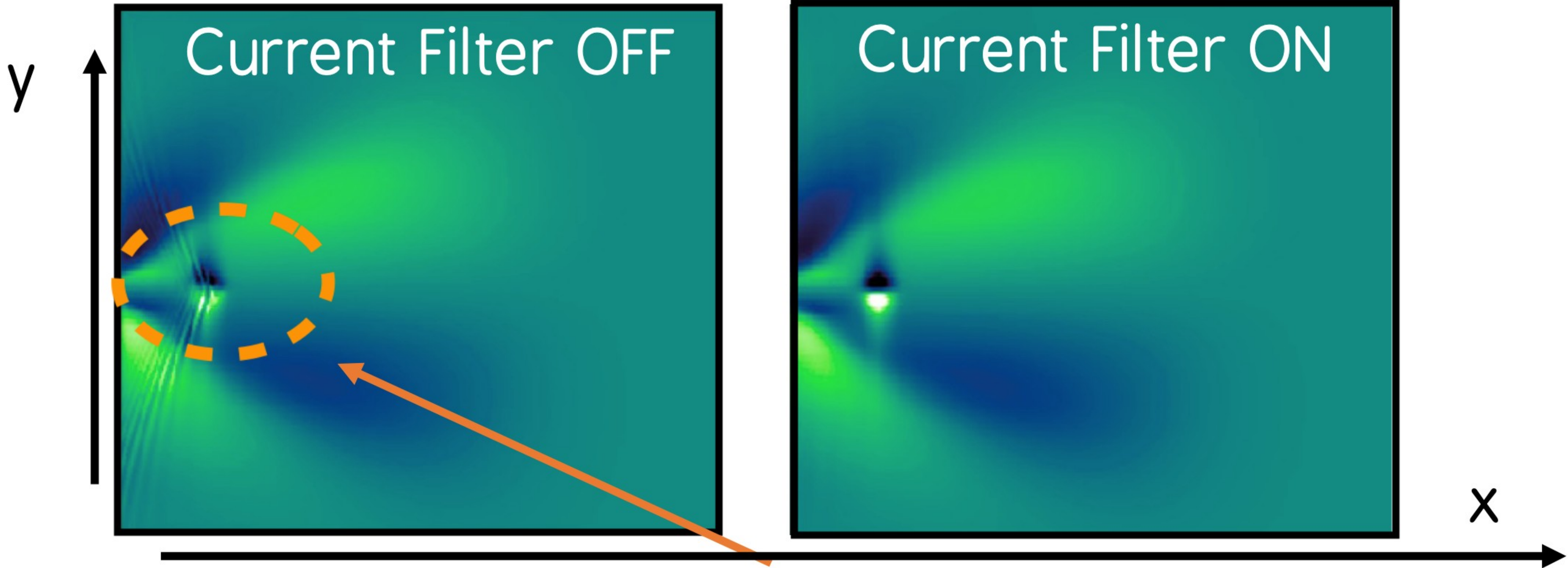
Write Maxwell's equations on a grid, and search for wave-like solutions

→ **numerical dispersion relation** $\Delta t^{-2} \sin^2(\omega\Delta t/2) = \sum_{a=x,y,z} \Delta a^{-2} \sin^2(k_a\Delta a/2)$



The numerical vacuum is dispersive and anisotropic!

Dispersive vacuum \rightarrow numerical Cherenkov



Filtering can reduce Numerical Cherenkov Radiation

The time step cannot be too large

From the dispersion relation, we can calculate the phase velocity.

Stability requires $v_{ph}\Delta t < \text{cell size}$

(i.e. light waves cannot cross 2 cells in 1 time step)

$$\rightarrow \Delta t^{-2} > \sum_{a=x,y,z} \Delta a^{-2}$$

Courant-Friedrich-Levy (CFL) condition

The cell size cannot be too large either

Depending on the situation you may need to resolve:

- ✓ The Debye length (or the simulation will have numerical heating)
- ✓ The laser wavelength (or it won't propagate)
- ✓ The skin depth



Often, a PIC simulation won't crash when the results are meaningless.

Users must understand the limitations and *test*.

An *Electromagnetic* PIC code:

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What physics are included? Or not?

PIC has ...

Vlasov → No interaction between individual particles

Maxwell → No quantum mechanics

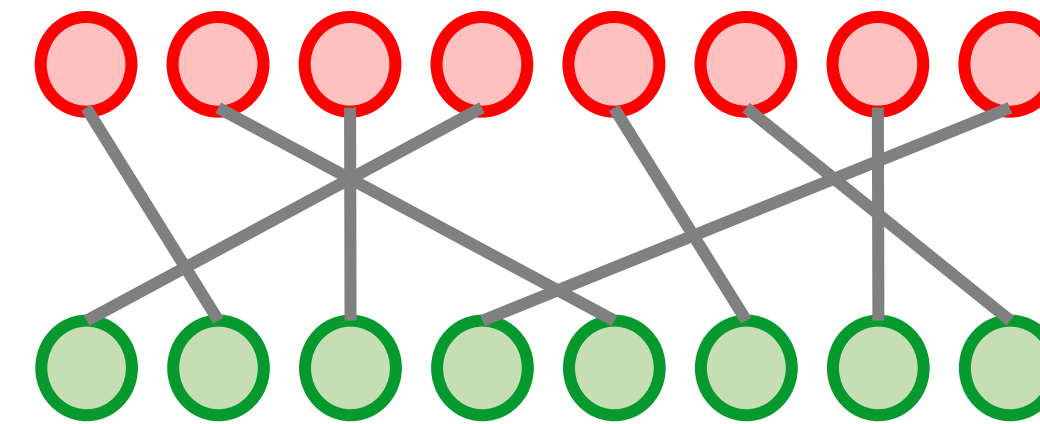
Fields on a grid → No high-frequency photons

Particles with fixed charge → No atomic physics

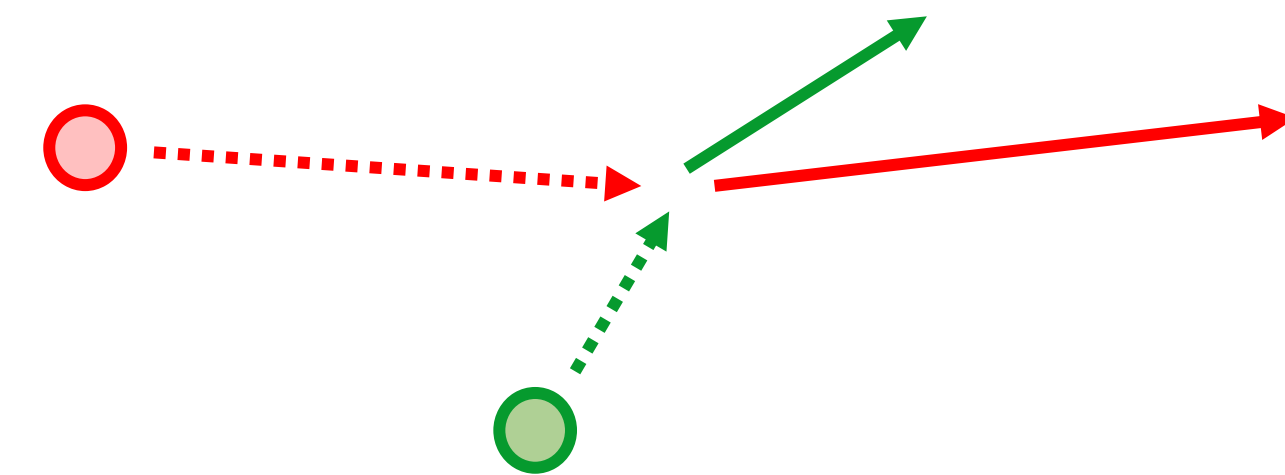
Particles with fixed mass → No nuclear physics

Collisions can be introduced with a Monte-Carlo method (other methods exist)

- Instead of calculating $n!$ interactions, quasi-particles are associated 2-by-2 randomly



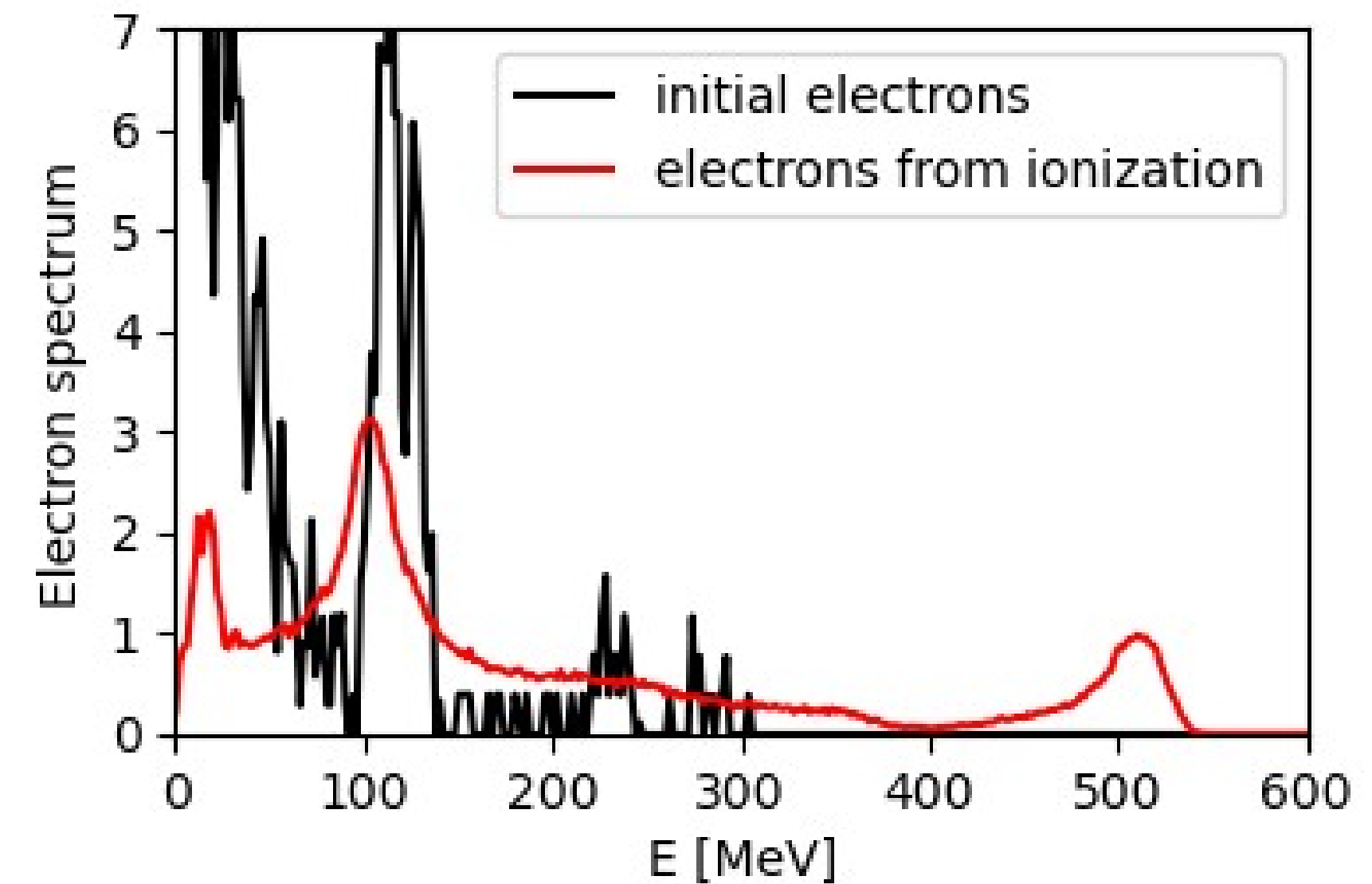
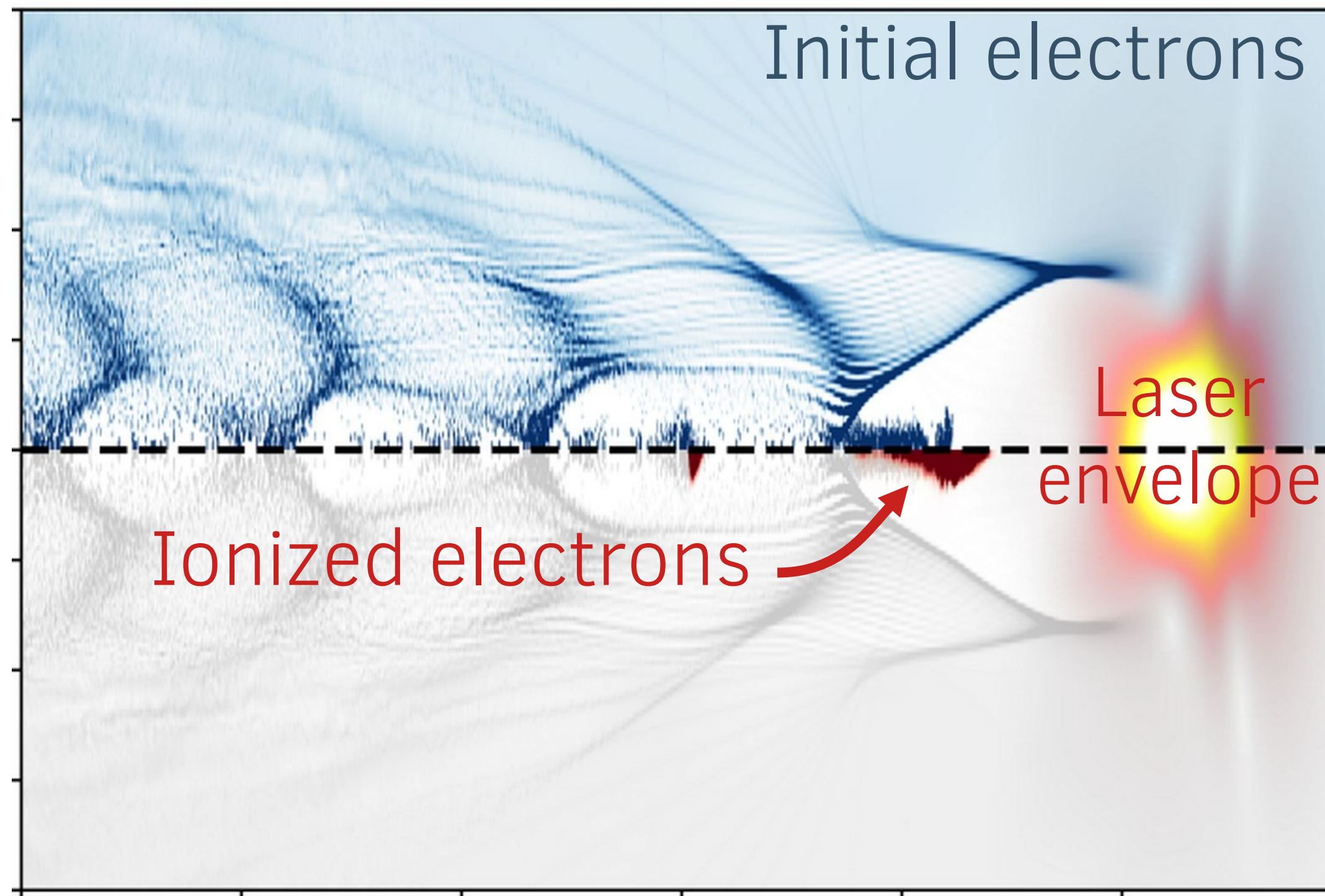
- The collision rate is computed for each pair
Usually: small-angle Rutherford cross-section.
- A random deflection is computed accordingly



→ provides additional physics:
stopping power, resistivity, heat transport, ...

Ionization by fields and by collisions may also be introduced with a Monte-Carlo method

Example with wakefield acceleration

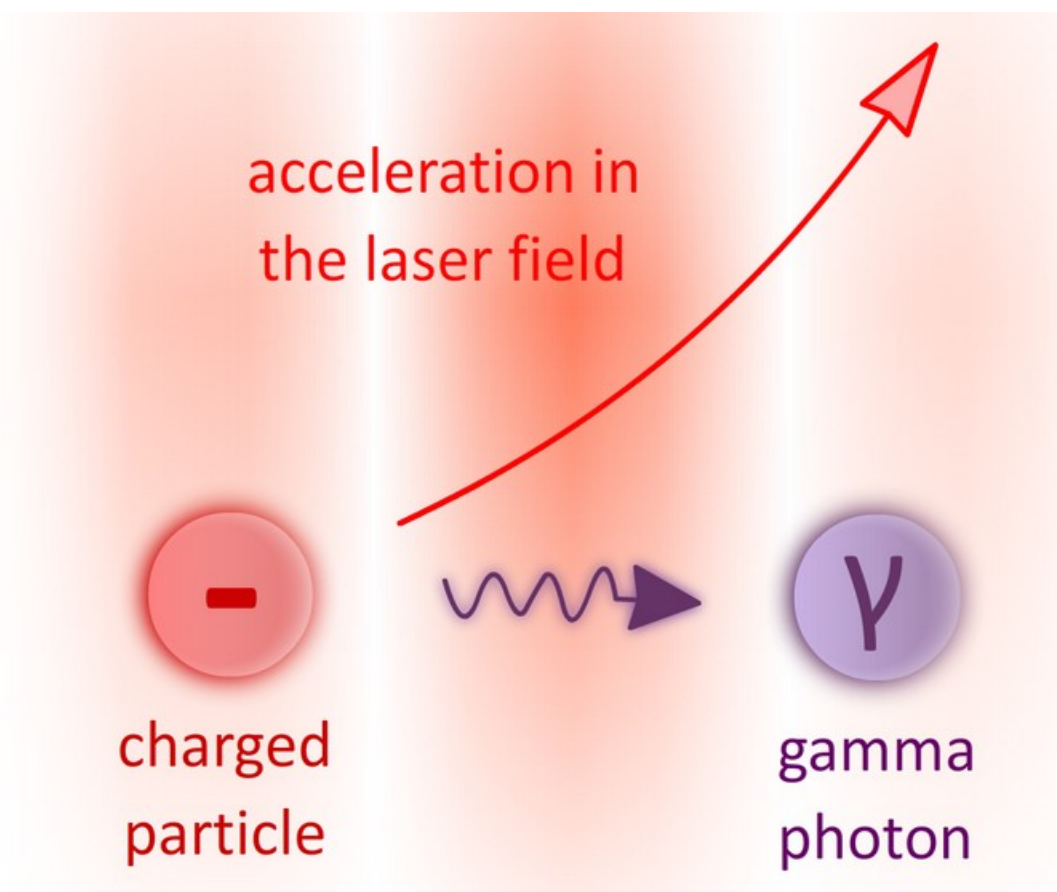


Strong-field quantum electrodynamics are relevant to multi-petawatt laser facilities

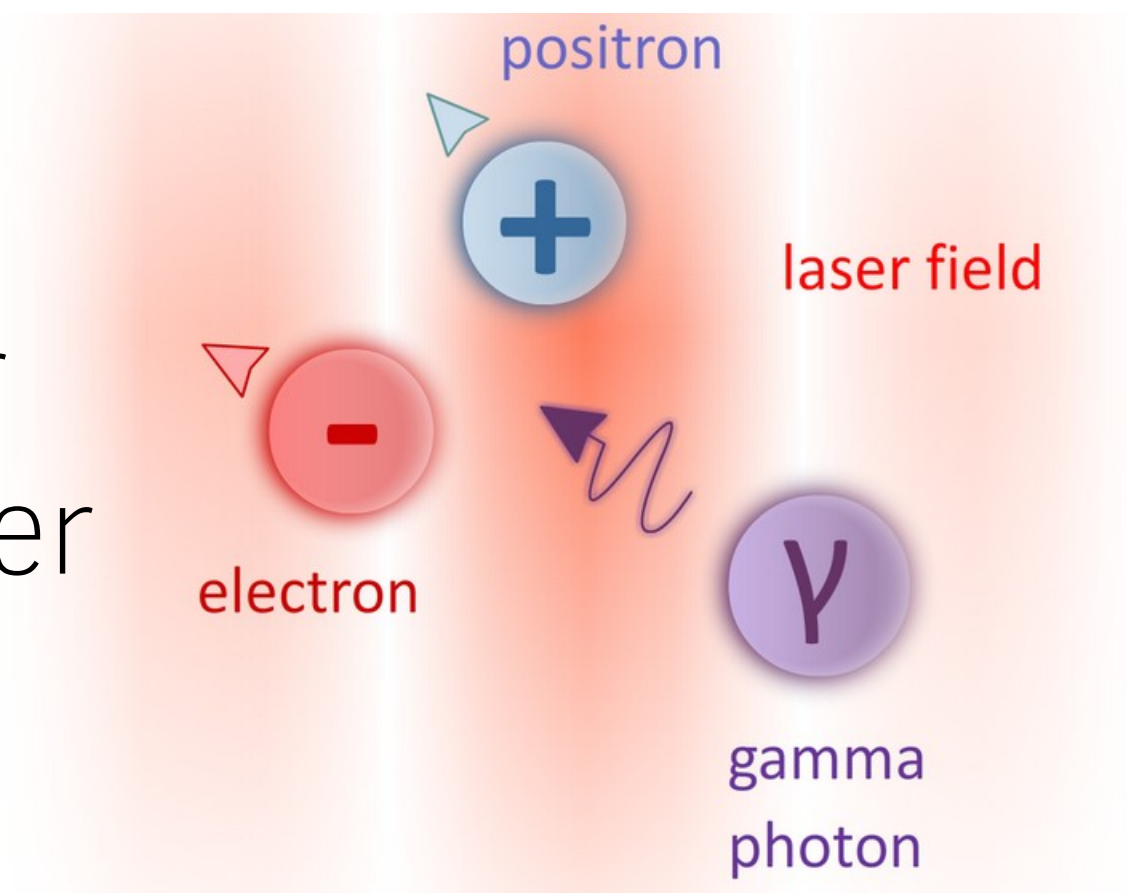
electron \rightarrow photon

photon \rightarrow e^-/e^+ pair

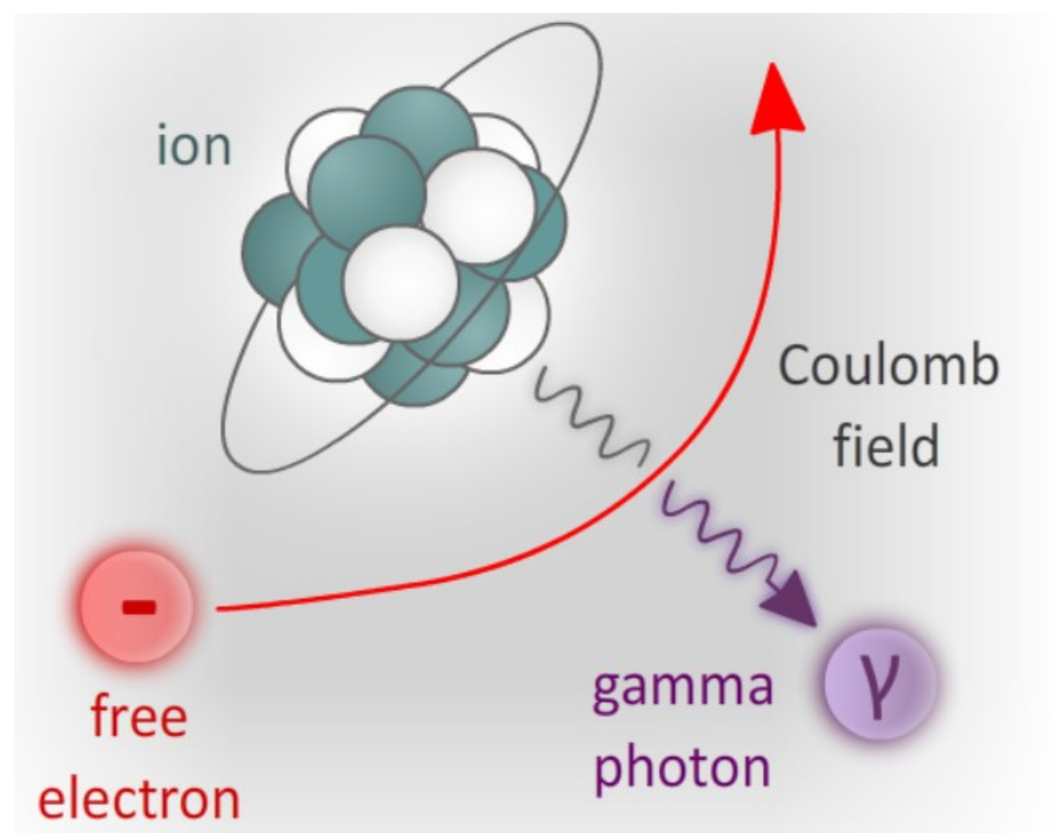
Non-linear Compton scattering



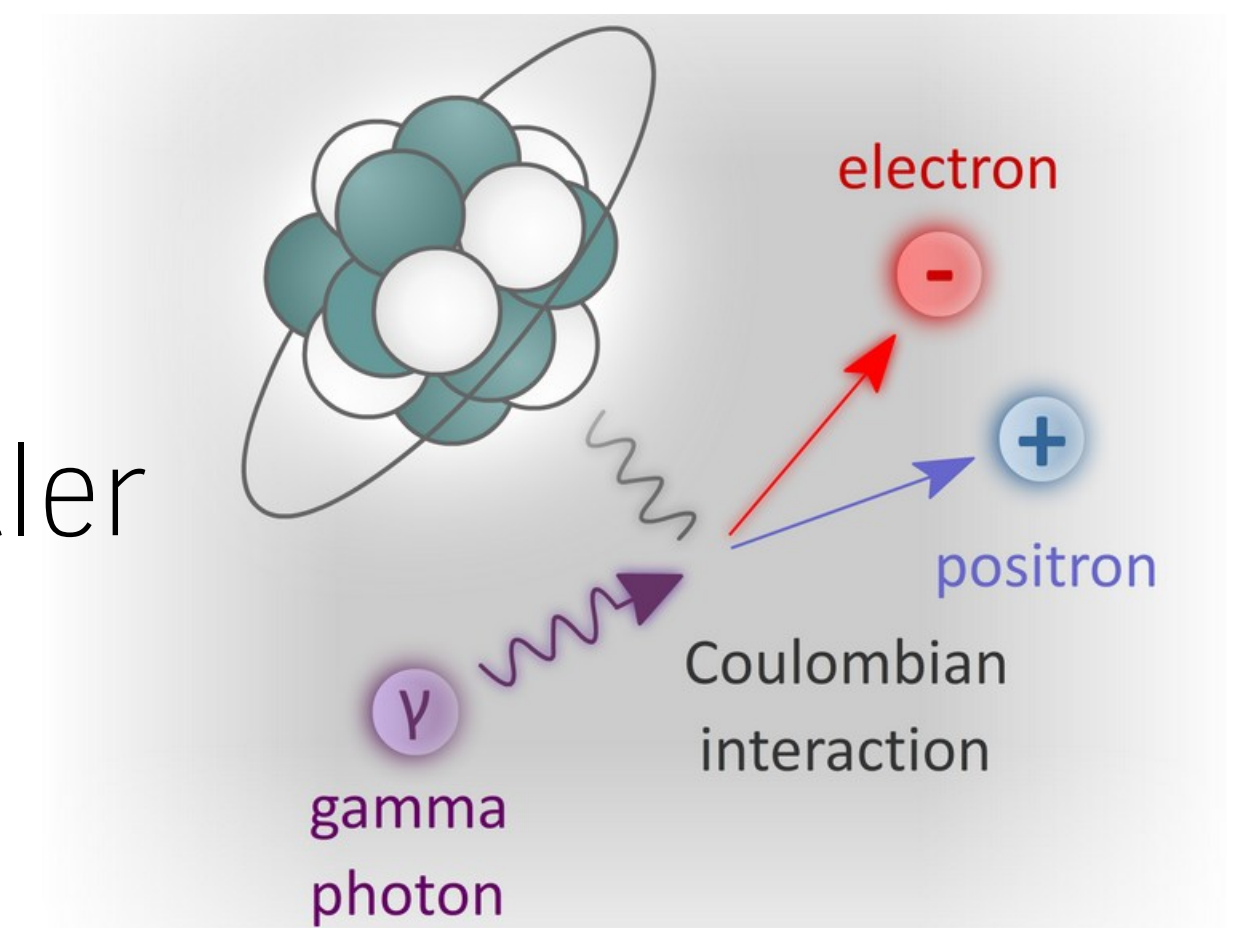
Non-linear Breit-Wheeler



Bremsstrahlung



Bethe-Heitler



Many other physics modules are possible

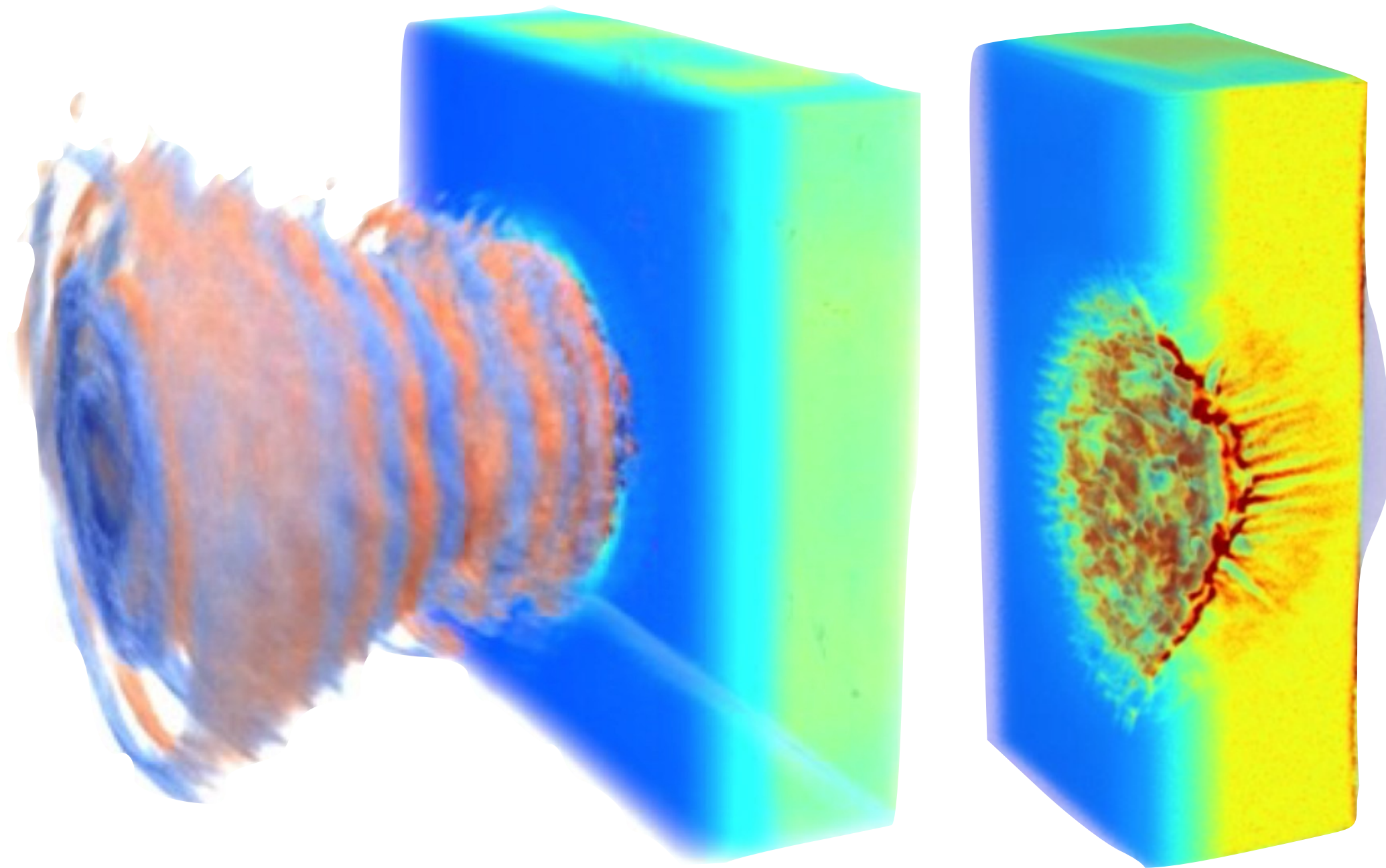
- Nuclear reactions
- Radiation transport
- Fluid species
- Particles injected from the boundary
- ...

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How long does a simulation take?

Example: Ultra-high intensity laser-solid interaction



High density's Debye length ~ 2 nm
target size $\sim 1 \mu\text{m} \times 10 \mu\text{m} \times 10 \mu\text{m}$

- 10^{10} cells
- 10^{11} quasi-particles
- 10^{13} bytes = **10 TeraBytes !**

Duration 400 fs, maximum timestep ~ 0.004 fs

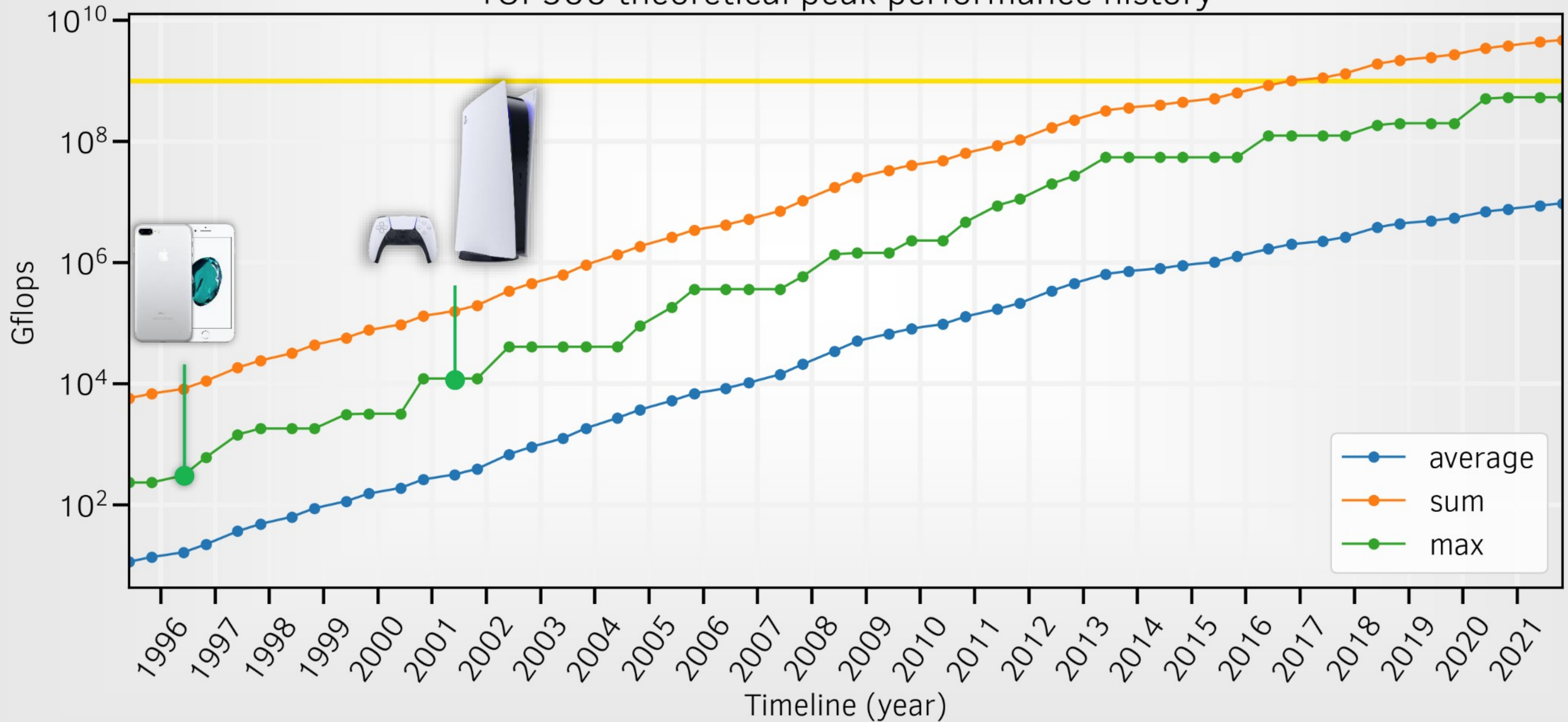
- 100000 time steps
- 10^7 operations/particle
- 10^{18} operations in total

1 CPU @ 2 GHz → **16 years !**

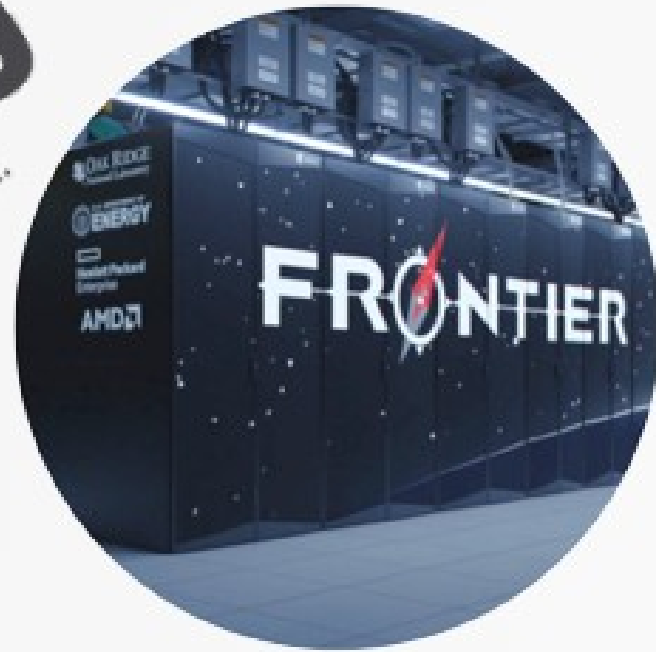
Solution: use 100 000 CPUs → 12h simulation

Supercomputers grow exponentially in performance

TOP500 theoretical peak performance history



The exascale challenge



Frontier (US), **1100 Pflops**
8.7 Mcores, **21 MW**
AMD CPU + accel. (GPU)



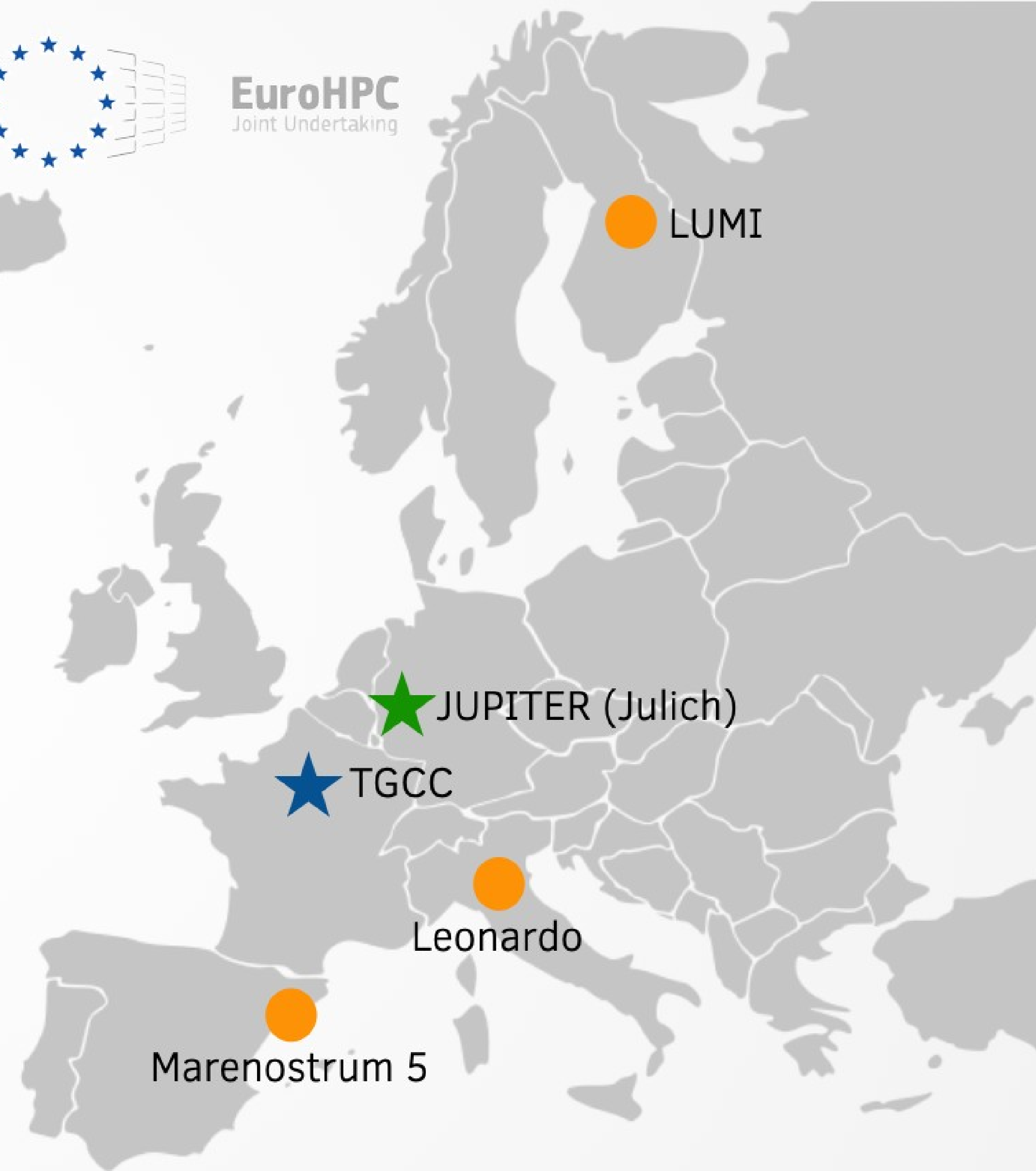
Fugaku (Japan), **442 Pflops**
7.6 Mcores, **21 MW**
Fujitsu ARM A64X



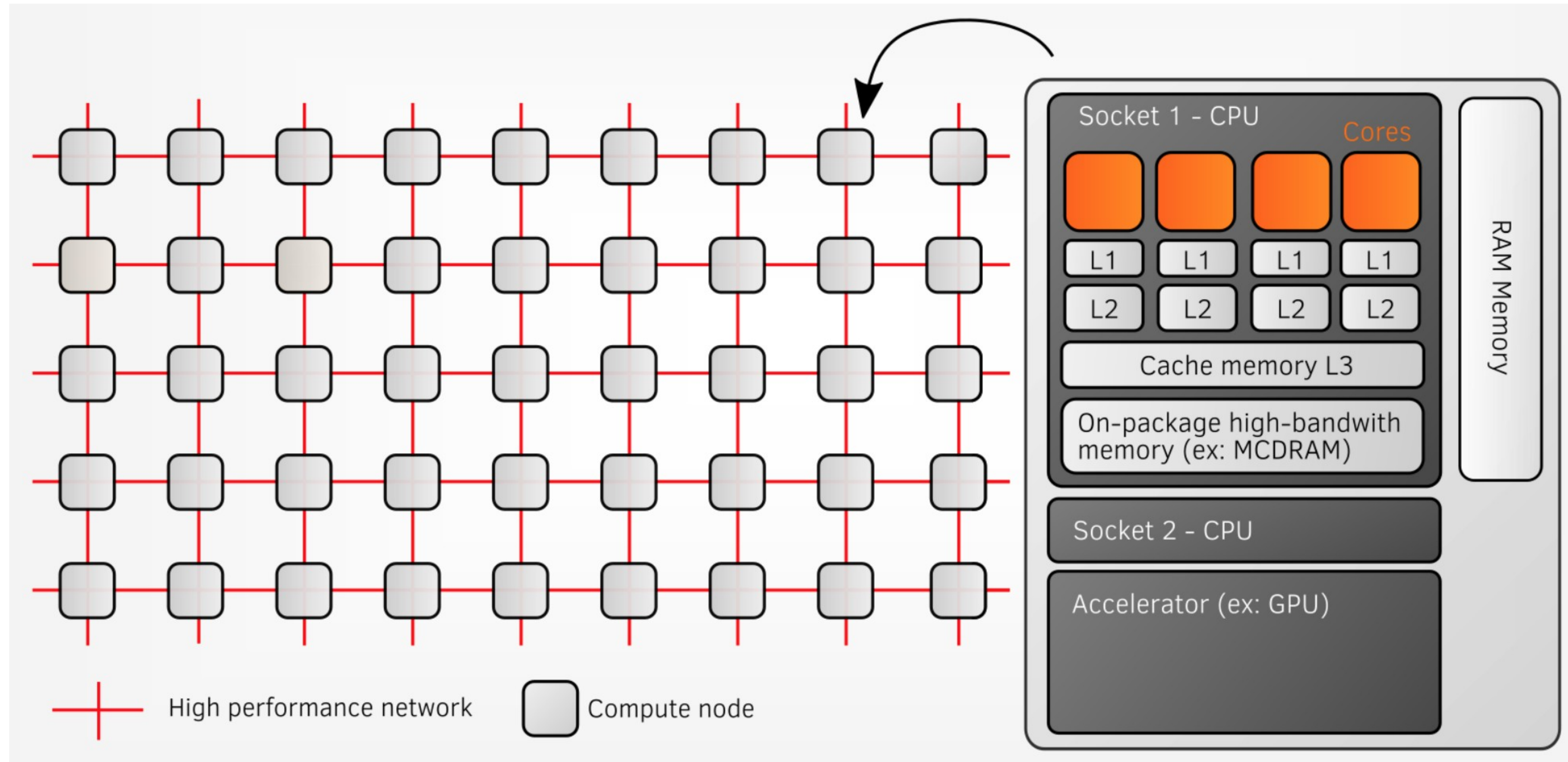
LUMI (Finland, EuroHPC), **152 Pflops**
1.1 Mcores, **3 MW**
AMD CPU + accel. (GPU)



EuroHPC
Joint Undertaking



Supercomputers are complex machines



It is extremely challenging to use efficiently this kind of computer

PIC users must know
how to carefully setup supercomputers

- Choose the decomposition of the simulation box
- Distribute the workload across CPUs / GPUs
- Test many possible configurations

Summary

PIC codes are popular and can address **many plasma physics problems**

The electromagnetic PIC method **solves the Vlasov+Maxwell system**

It has important **limitations and stability requirements**

Other physics can be added (collisions, ionization, QED, ...)

Supercomputers are often needed, and only modern codes can tackle recent development in computer architectures

Thank you for your attention

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ELI Summer School 2023

Smilei)

